



Fermi

Gamma-ray Space Telescope



FGST Science Highlights

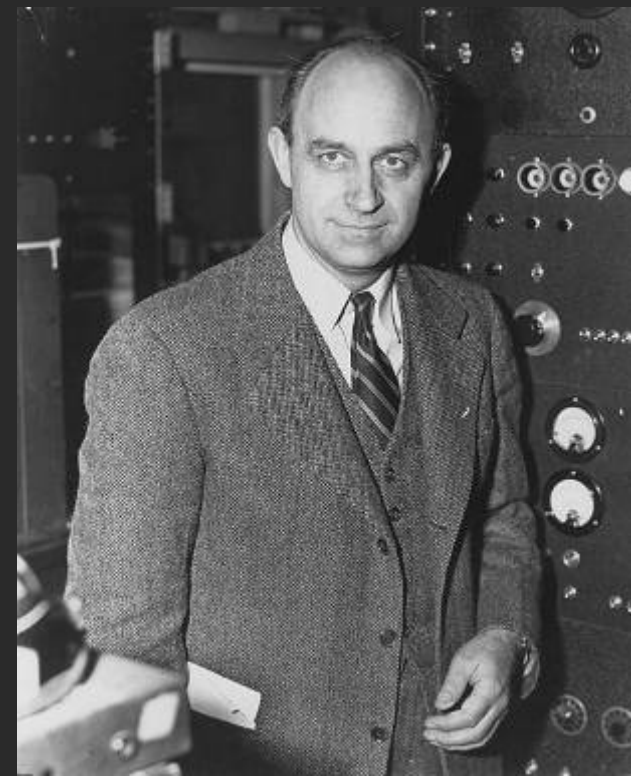
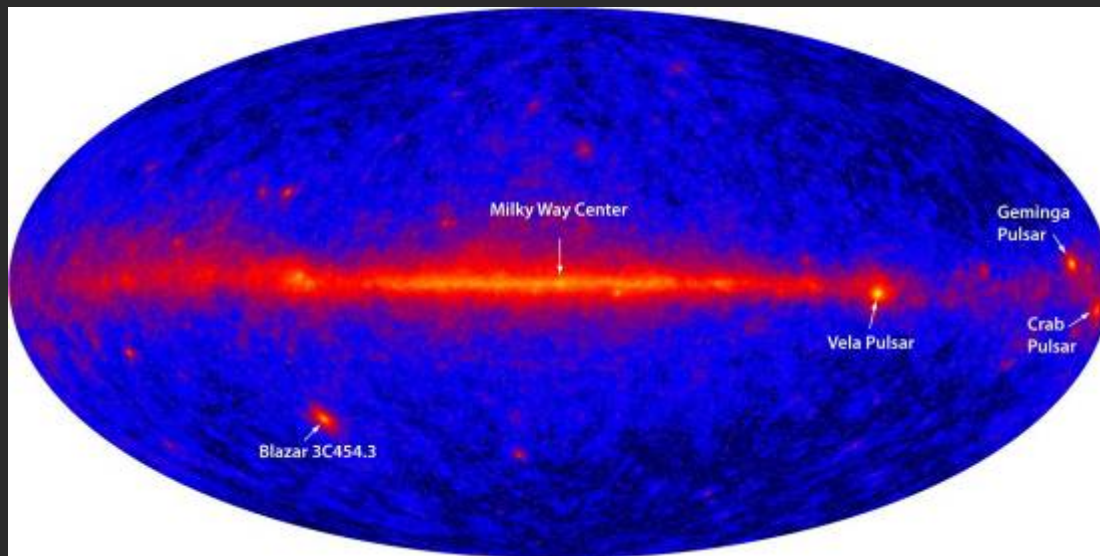
Launched 2008 June 11



Andreas von Kienlin (MPE)

First light and Observatory Renaming

- ◆ GLAST becomes Fermi Gamma-ray Space Telescope
 - “first light” 4 day sky map



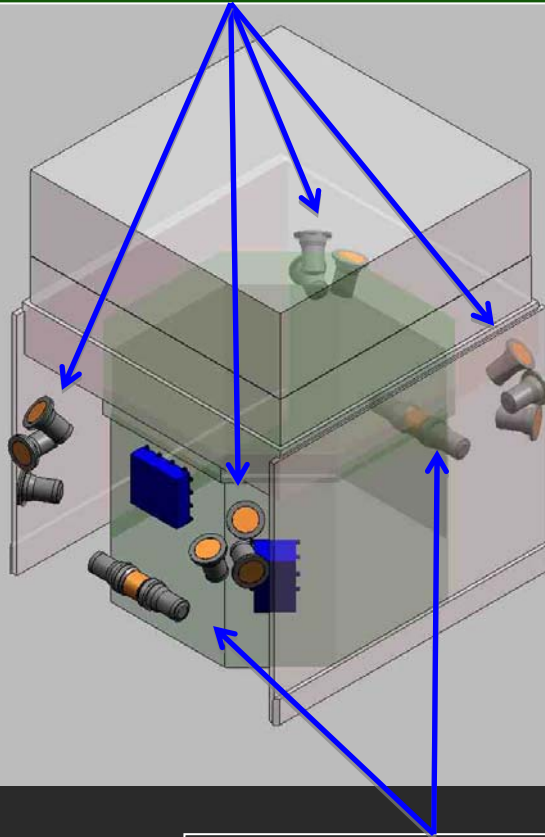
“ Enrico Fermi (1901-1954) was an Italian physicist who immigrated to the United States. He was the first to suggest a viable mechanism for astrophysical particle acceleration. This work is the foundation for our understanding of many types of sources to be studied by NASA’s Fermi Gamma-ray Space Telescope, formerly known as GLAST. ”

The Fermi Observatory

Large Area Telescope (LAT)

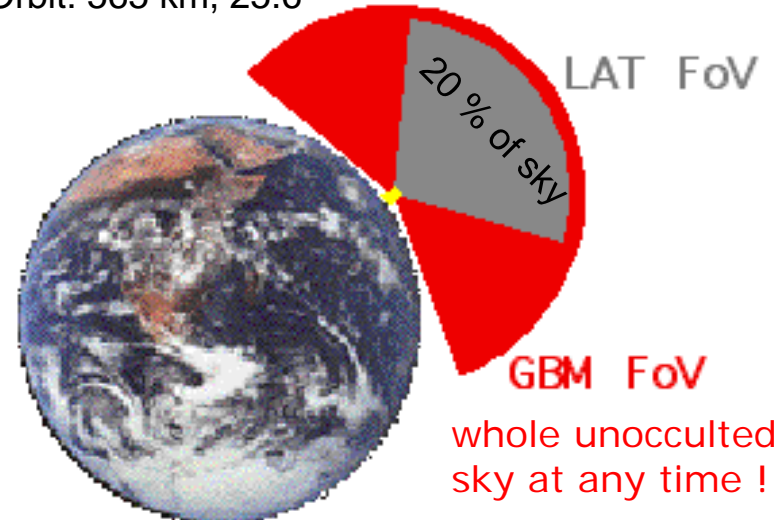


Gamma-ray Burst Monitor (GBM)



GBM BGO detectors

Orbit: 565 km, 25.6°

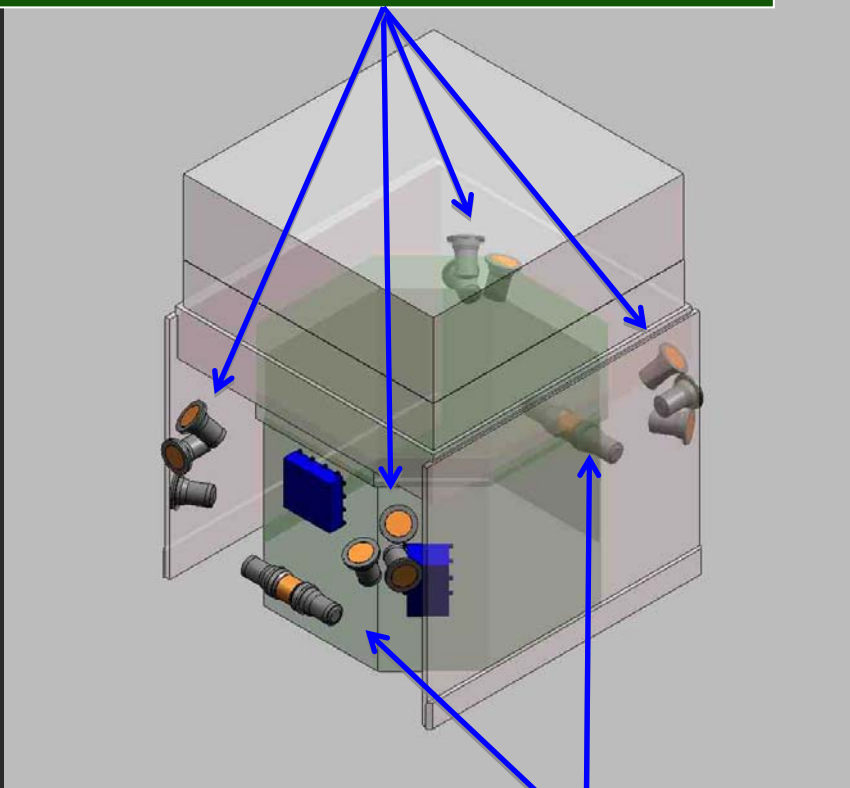


The Fermi Observatory

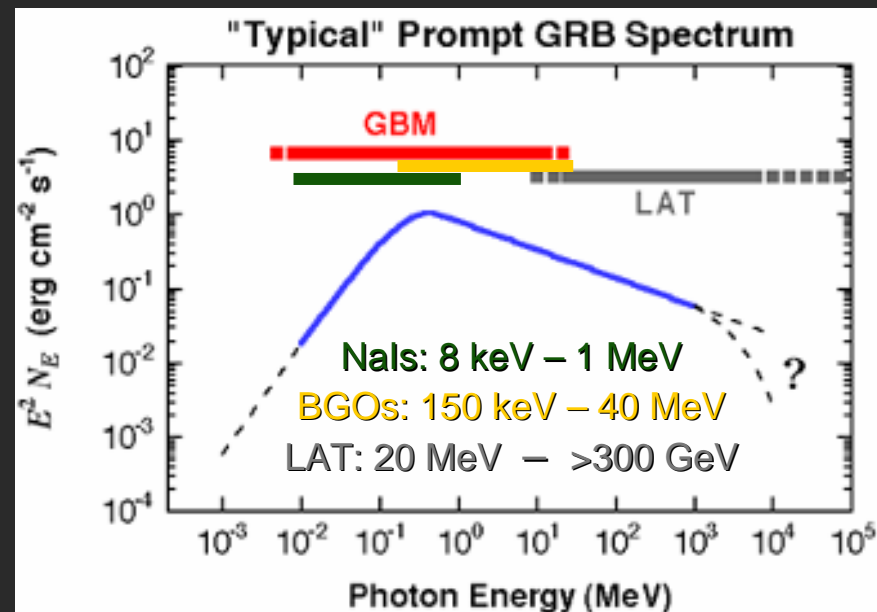
LAT (high-E spectrum)



NaIs (location & low-E spectrum)



BGOs (mid-E spectrum)



Fermi Gamma-Ray Burst Monitor

12 × Sodium Iodide NaI(TL) scintillation detectors

⇒ build by MPE + MSFC/UAH

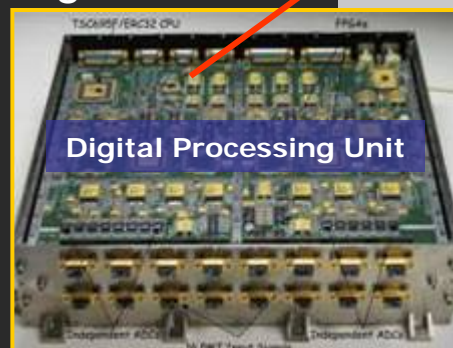
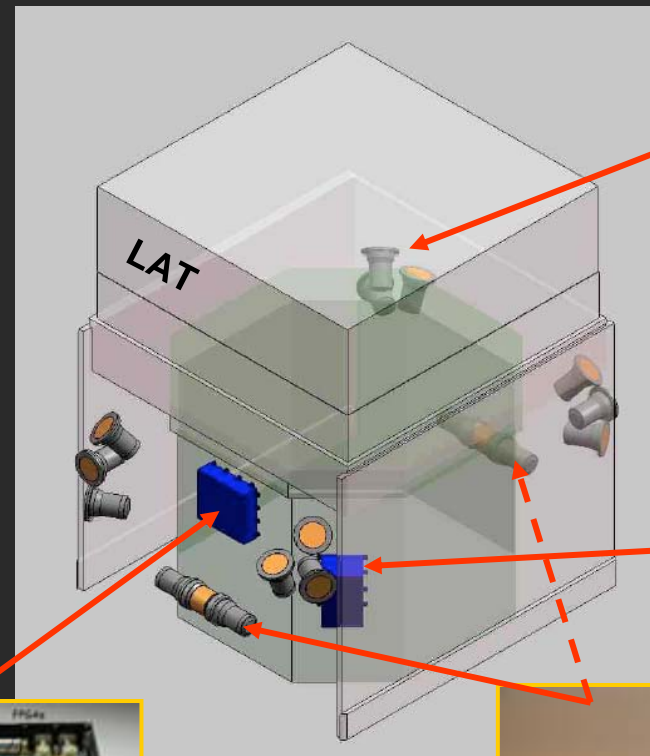
- Wide Field of View
- Burst Trigger
- Coverage of typical GRB spectrum: **8 keV – 1 MeV**

2 × Bismuth Germanate (BGO) scintillation detectors

- Spectral overlap with the LAT: **150 keV-40 MeV**

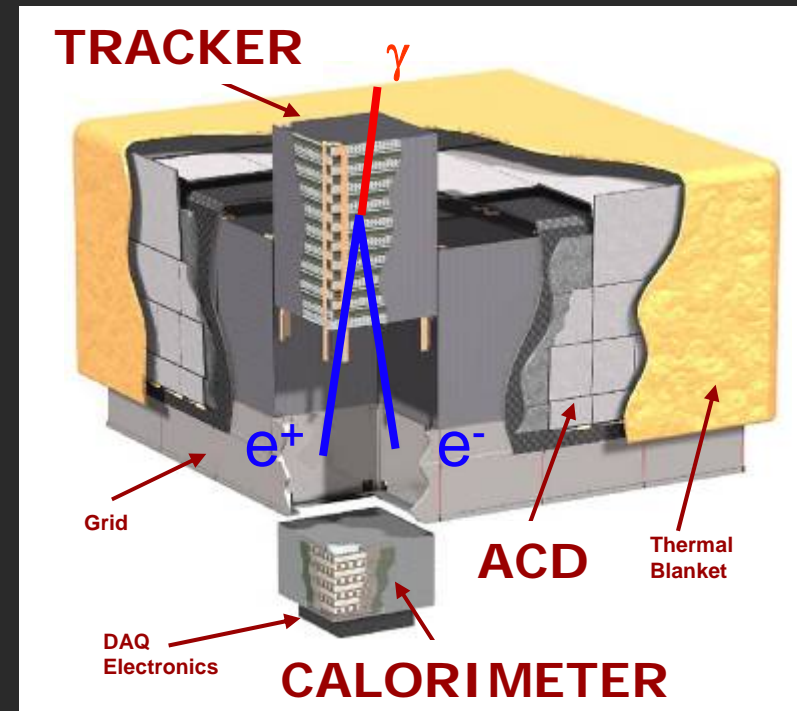
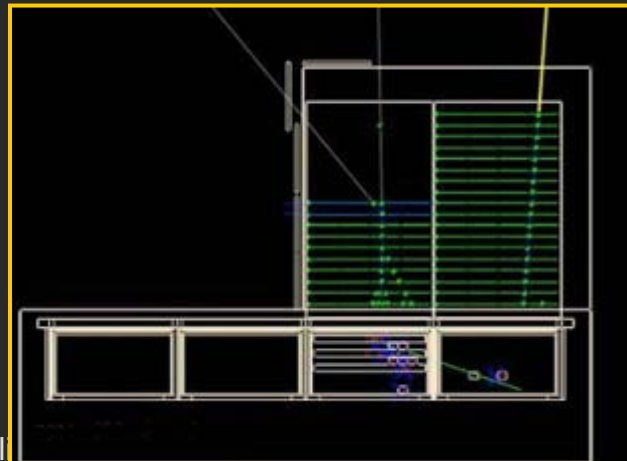
1 × Power Box (PB)

1 × Digital Processing Unit (DPU)



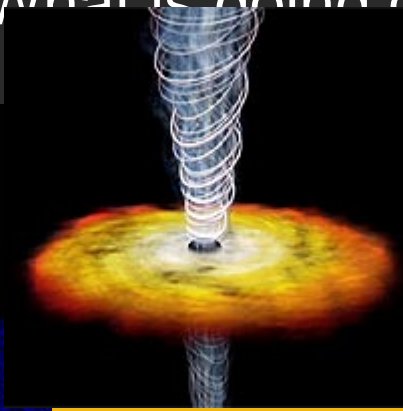
Fermi Large-Area Telescope (LAT)

- ◆ Pair Conversion
 - 4 × 4 modular array, 3000 kg, 650 W
 - Detect photons between ~20 MeV - 300GeV
- ◆ Tracking system
 - Silicon-Strip Detectors (880000 channels)
- ◆ Calorimeter
 - CsI Crystals (8.4 r.l., hodoscopic array)
- ◆ Anticoincidence
 - Segmented ACD veto counters (tiles)



The Fermi Sky 100 MeV to 300 GeV

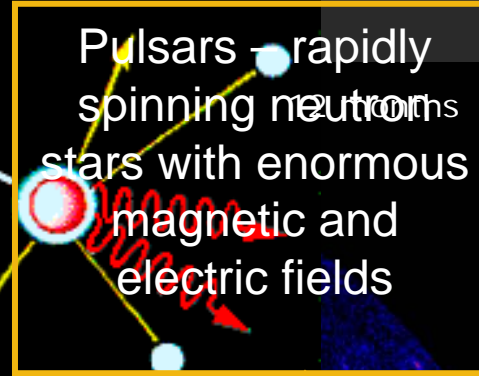
- ◆ What is going on in the gamma-ray sky?



Gamma rays
cosmic ray
shining into
us between
stars



Particle-Particle
Collisions

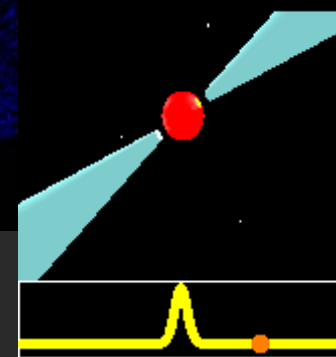


Pulsars – rapidly
spinning neutron
stars with enormous
magnetic and
electric fields

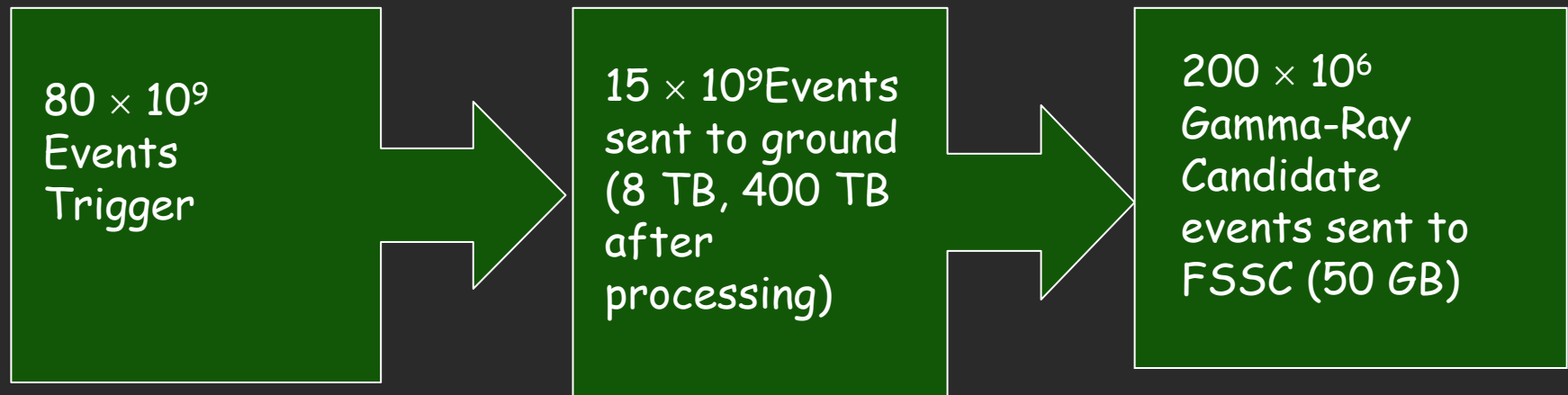


Blazars –
supermassive
black holes with
huge jets of
particles and
radiation pointed
right at Earth.

MPIfR-Bonn Pulsar Group

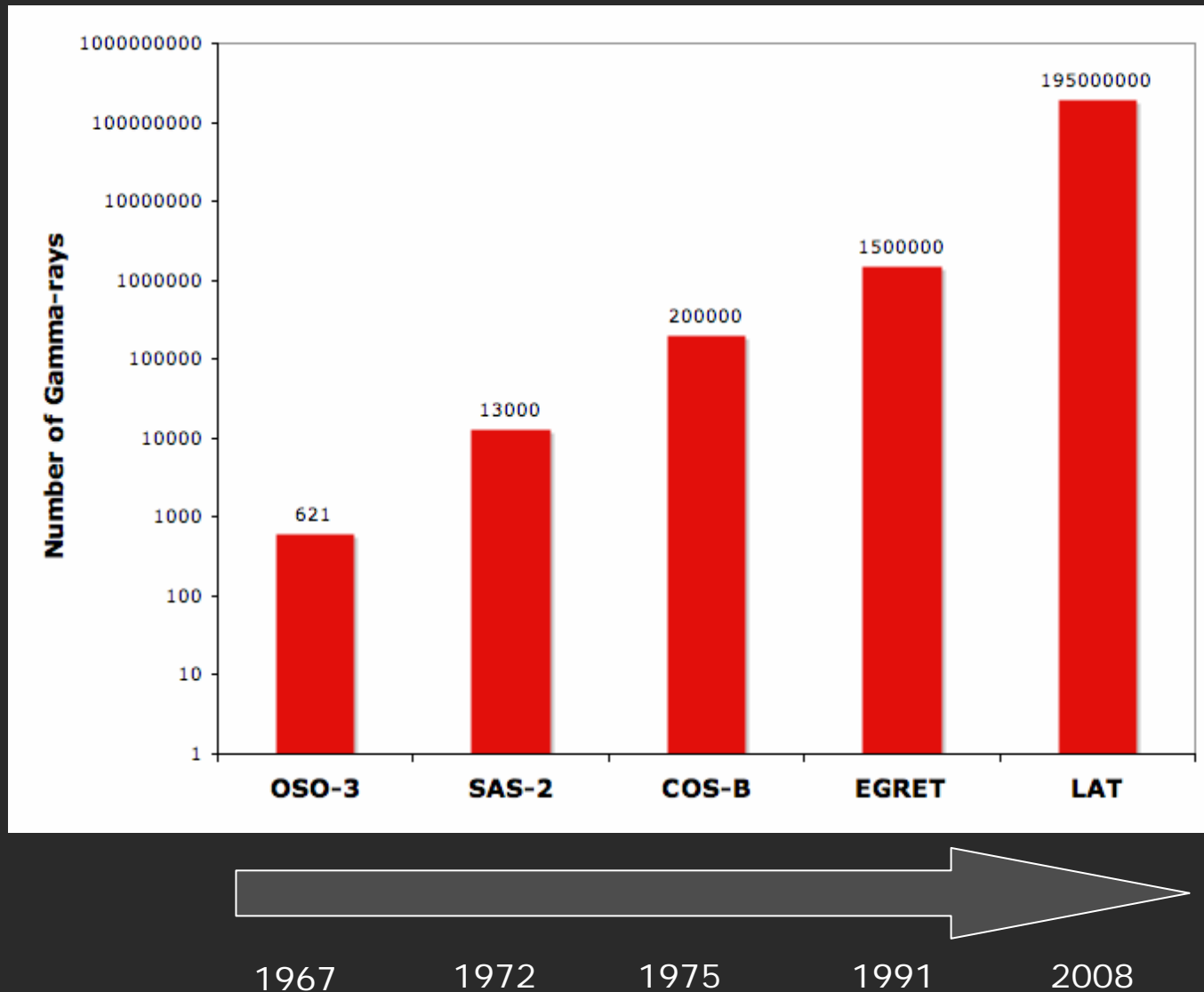


Fermi/LAT: data volume & processing

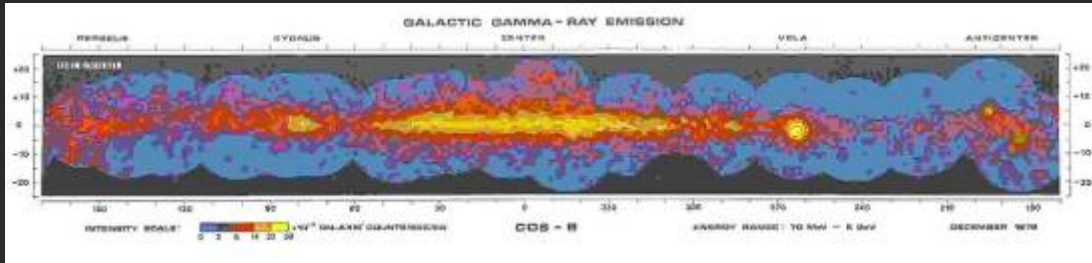


⇒ 160 CPU years in the last 16 months

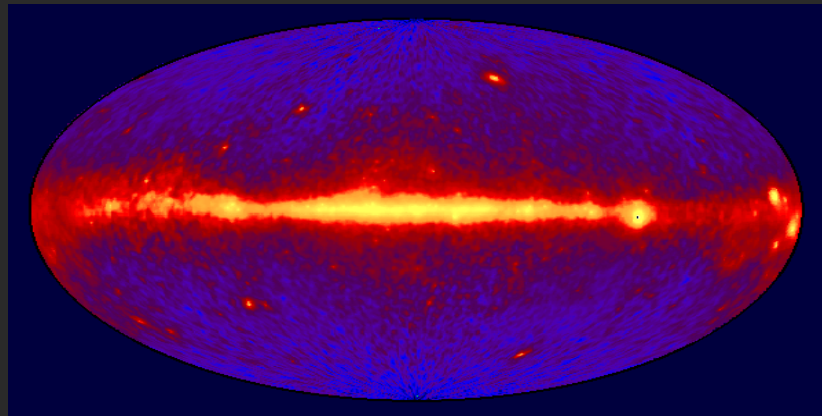
Number of Gamma-photons



The Milky Way at $E_\gamma > 100$ MeV

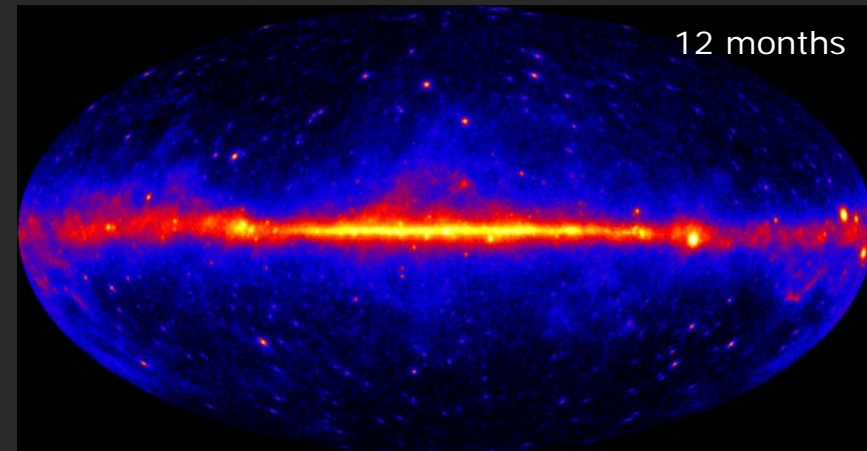


COS-B: 1975-82



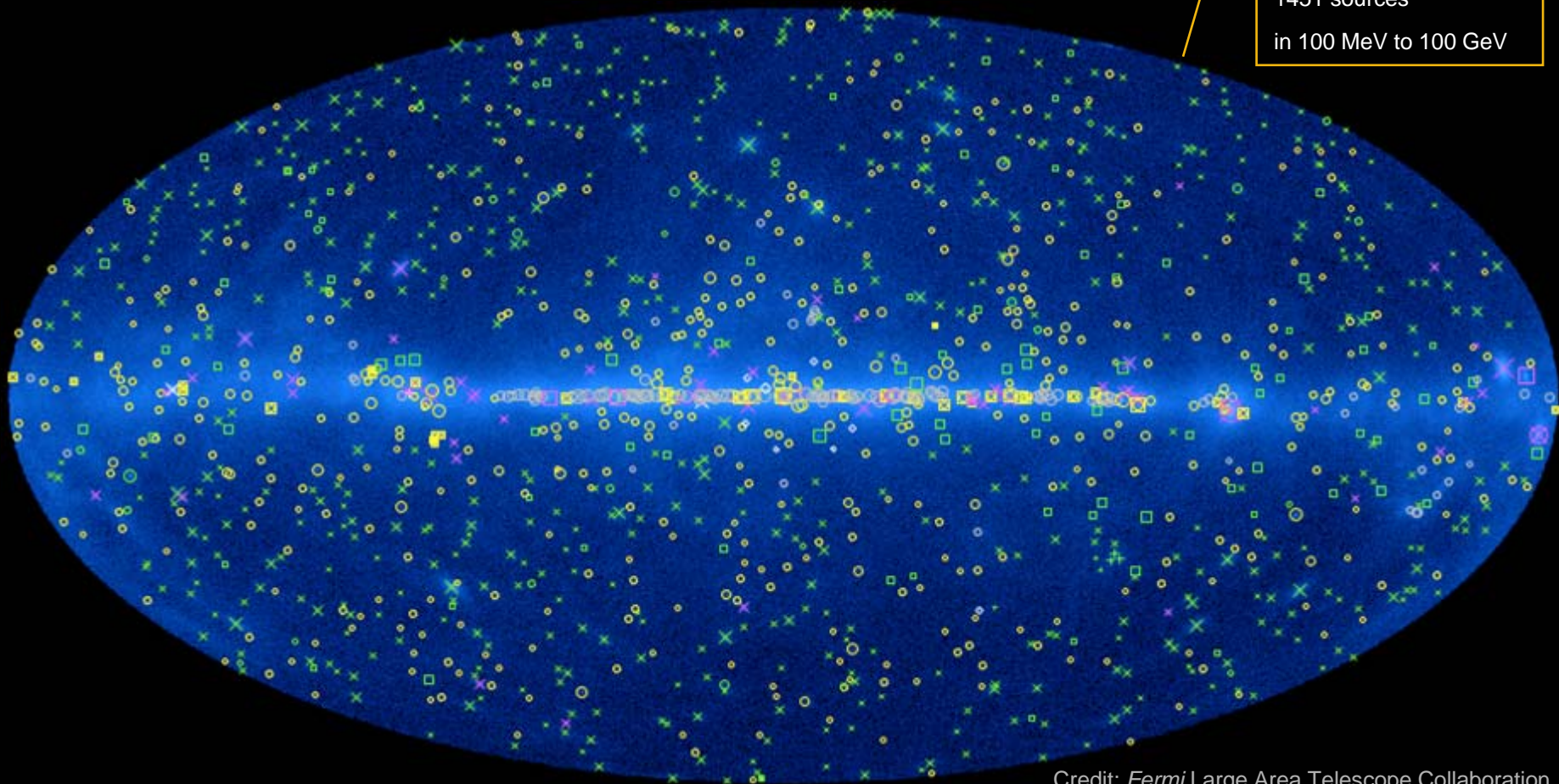
EGRET: 1991-2000

Fermi: Juli 2009



The Fermi LAT 1FGL Source Catalog

First 11 months:
1451 sources
in 100 MeV to 100 GeV



Credit: *Fermi* Large Area Telescope Collaboration

- | | | | | | | | |
|---|--|---|-------------------------|---|-------------------------|---|------------|
| ○ | AGN | □ | Starburst Galaxy | ⊗ | PSR w/PWN | □ | SNR |
| × | AGN-Blazar | + | Galaxy | ◇ | Globular Cluster | ○ | PWN |
| □ | AGN-Non Blazar | | | × | HXB or MQO | × | PSR |
| ○ | No Association | | | | | | |
| ◻ | Possible Association with SNR and PWN | | | | | | |
| ○ | Possible confusion with Galactic diffuse emission | | | | | | |

The First LAT AGN catalog (1LAC)

- ◆ 599 AGNs (+ ~100 candidates)

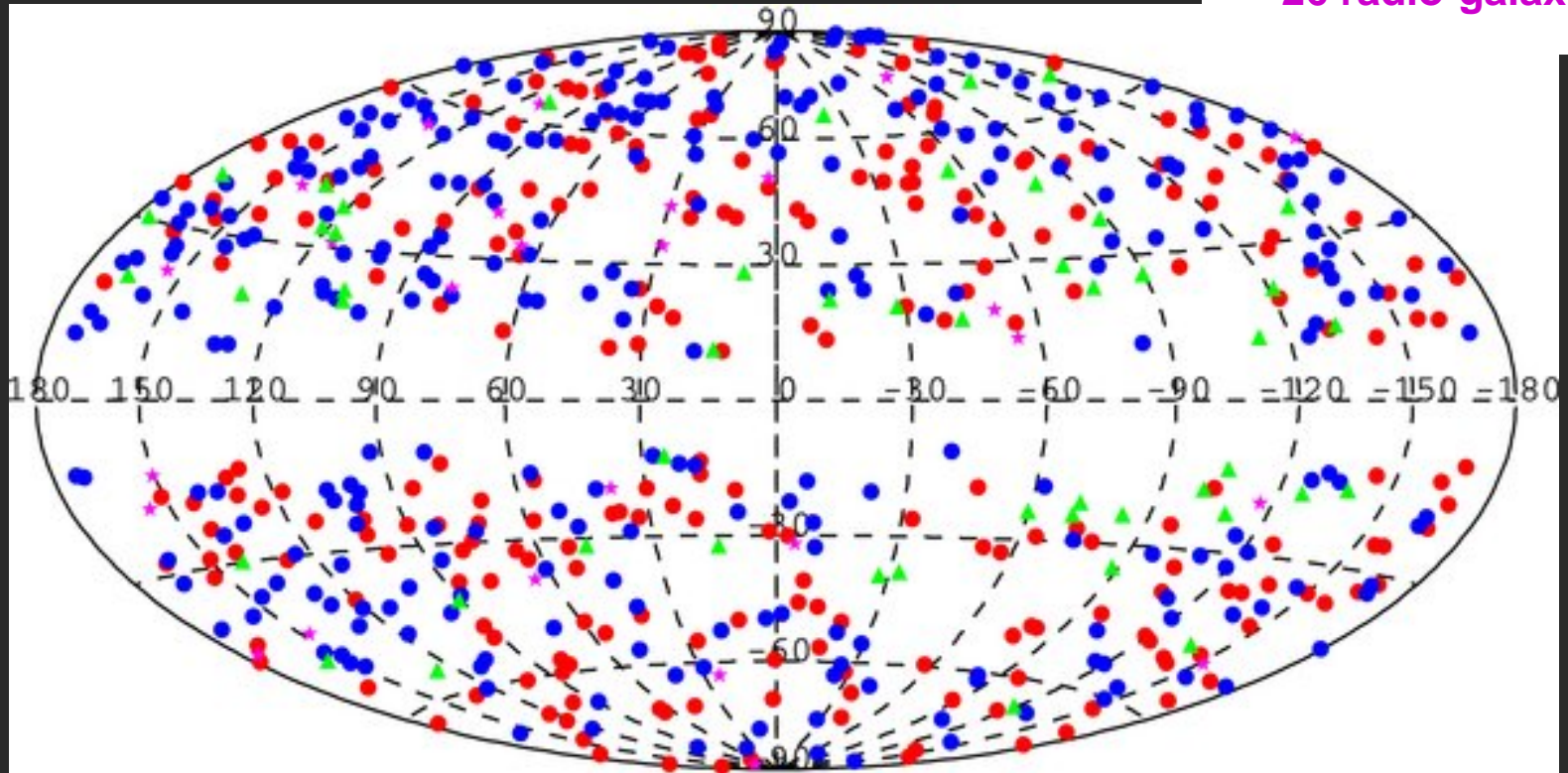
- 11 month data set (“clean sample”, $|b| > 10^\circ$)

248 FSRQs

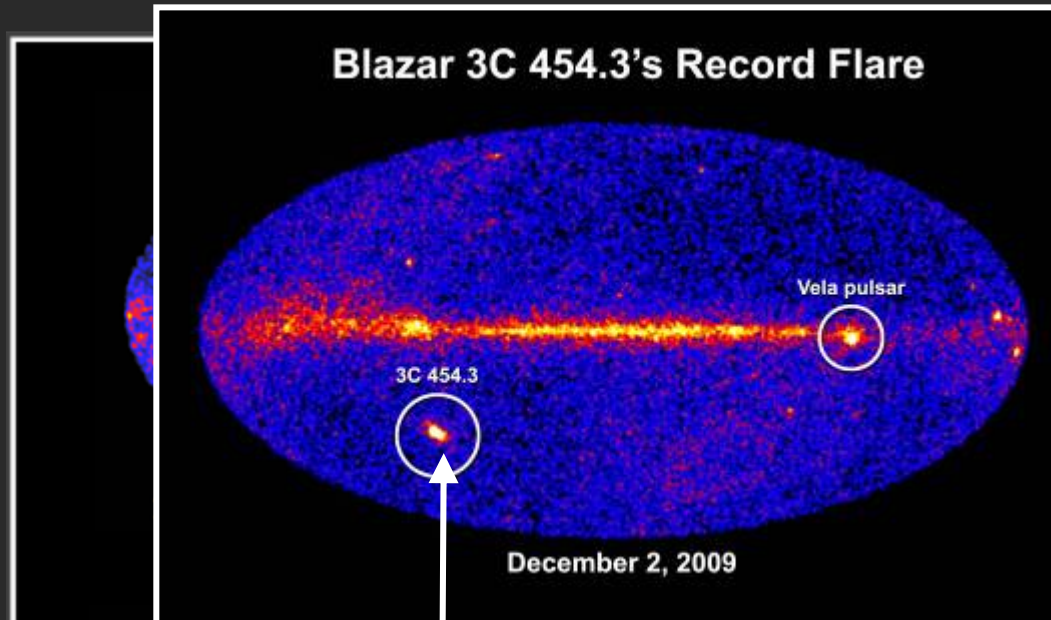
275 BLLacs

50 of unknow type

26 radio-galaxies



The flaring and variable sky



~50 Astronomers telegrams

◆ Discovery of **new gamma-ray blazars**:

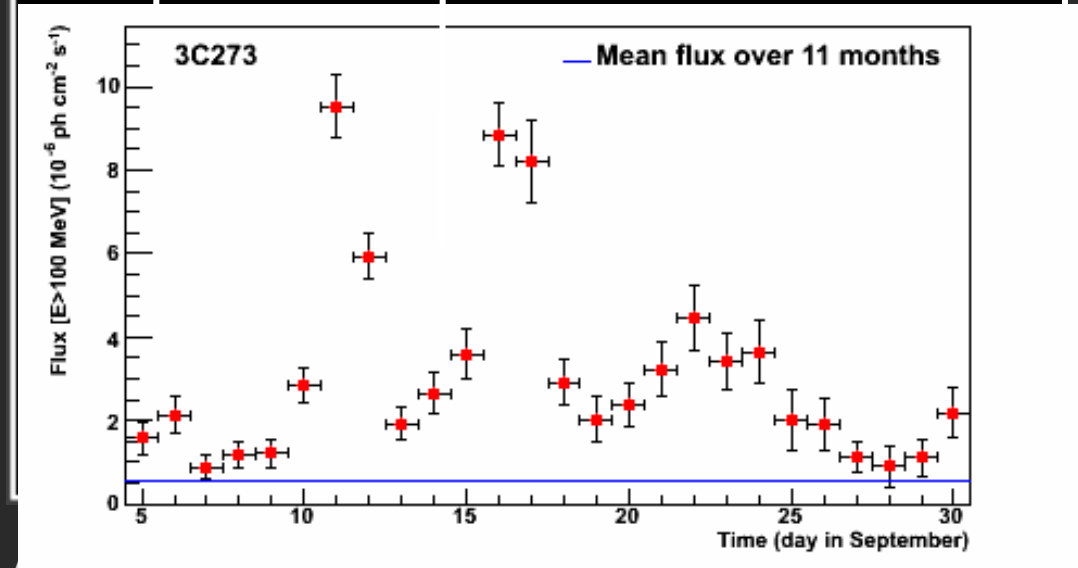
- PKS 1502+106,
- PKS 1454-354

◆ Flares from **known gamma-ray blazars**:

- 3C454.3, PKS 1510-089,
- 3C273, AO 0235+164,
- PSK 0208-512, 3C66A,
- PKS 0537-441.

◆ Galactic plane transients:

- J0910-5041,
- 3EG J0903-3531



Andreas von Kienlin

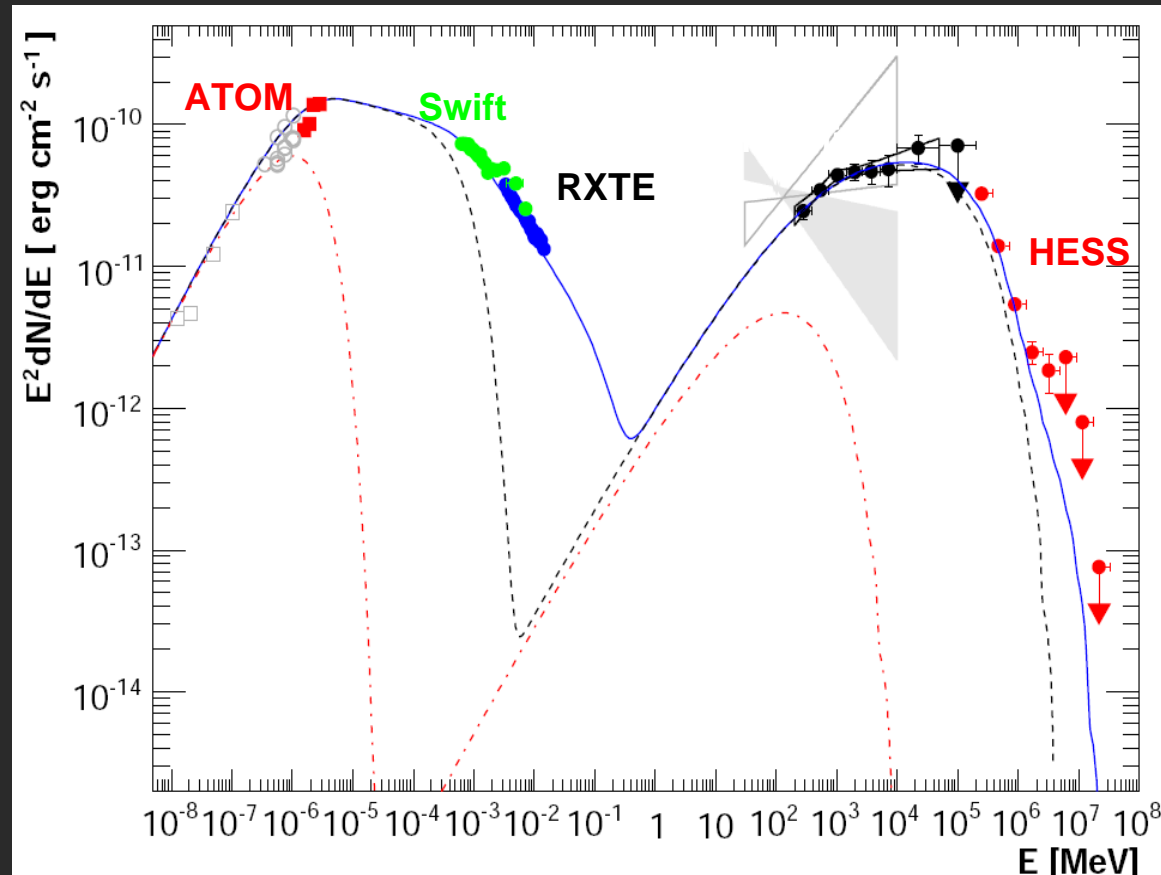
MW campaign on PKS 2155–304 (with HESS)

HSP-BLLac, $z=0.116$ nonflaring, low/quiescent state

First simultaneous SED including GeV-TeV

Unexpected correlations:

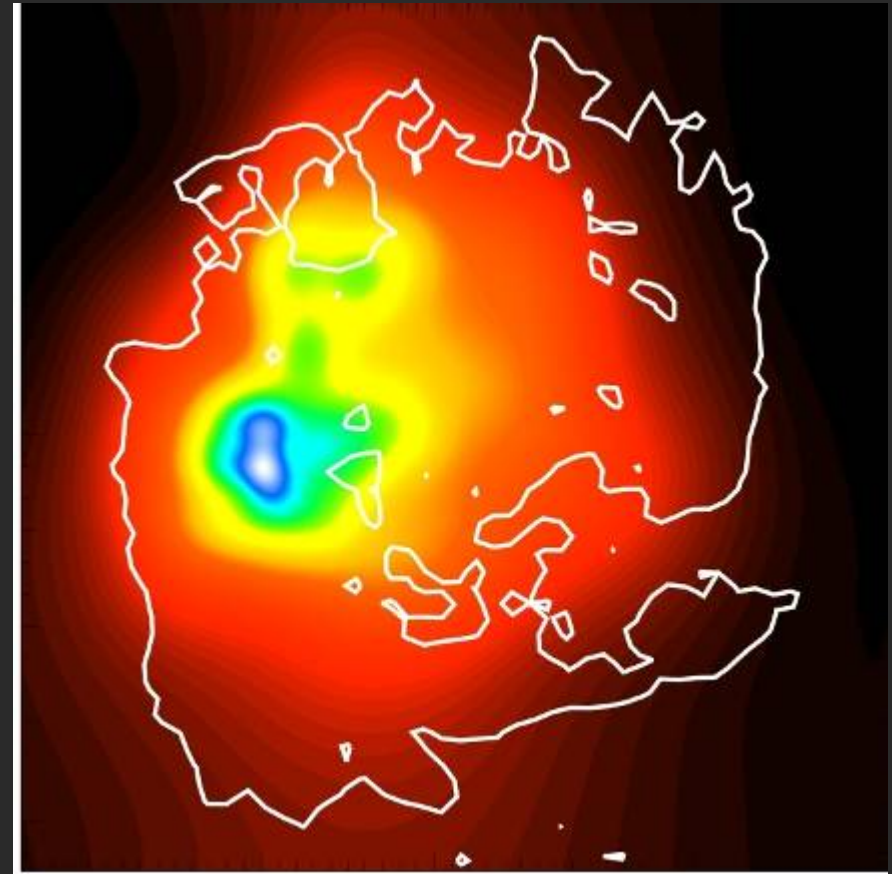
- strong correlation between optical and TeV fluxes
- X-ray flux varies independently of TeV flux
- correlation between X-ray flux and GeV photon index
- Challenge simple SSC models



Aharonian, F. et al. 2009, *ApJL*, 696 L150
contact authors: B. Giebels & J. Chiang

New category of Gamma-sources

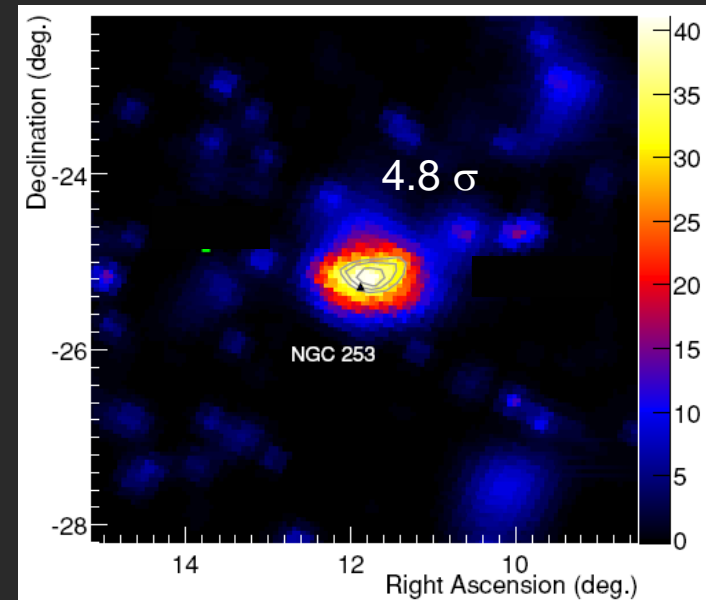
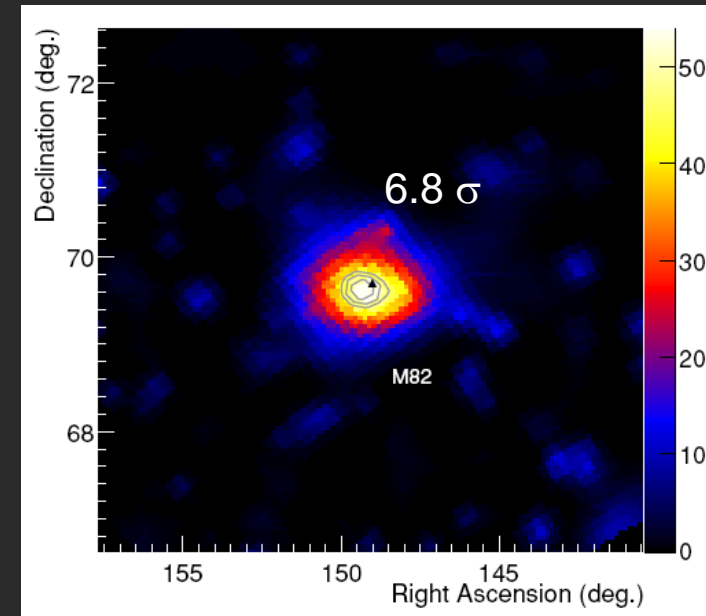
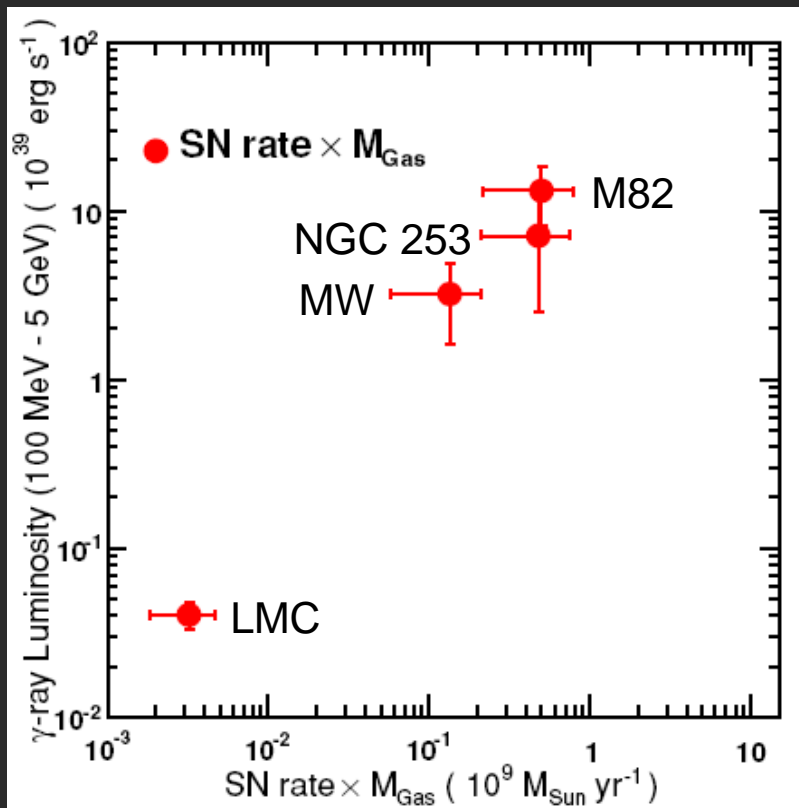
- ◆ Large Magellanic Cloud (LMC) for the first time resolved in gamma rays
 - 30 Doradus star forming region is a bright source of gamma rays and very likely a powerful cosmic-ray accelerator
 - Gamma-ray emission correlates well with massive star forming regions
 - CR interaction with interstellar medium and radiation field
 - Compactness of emission regions suggests little CR diffusion length



astro-ph, 1001.3298

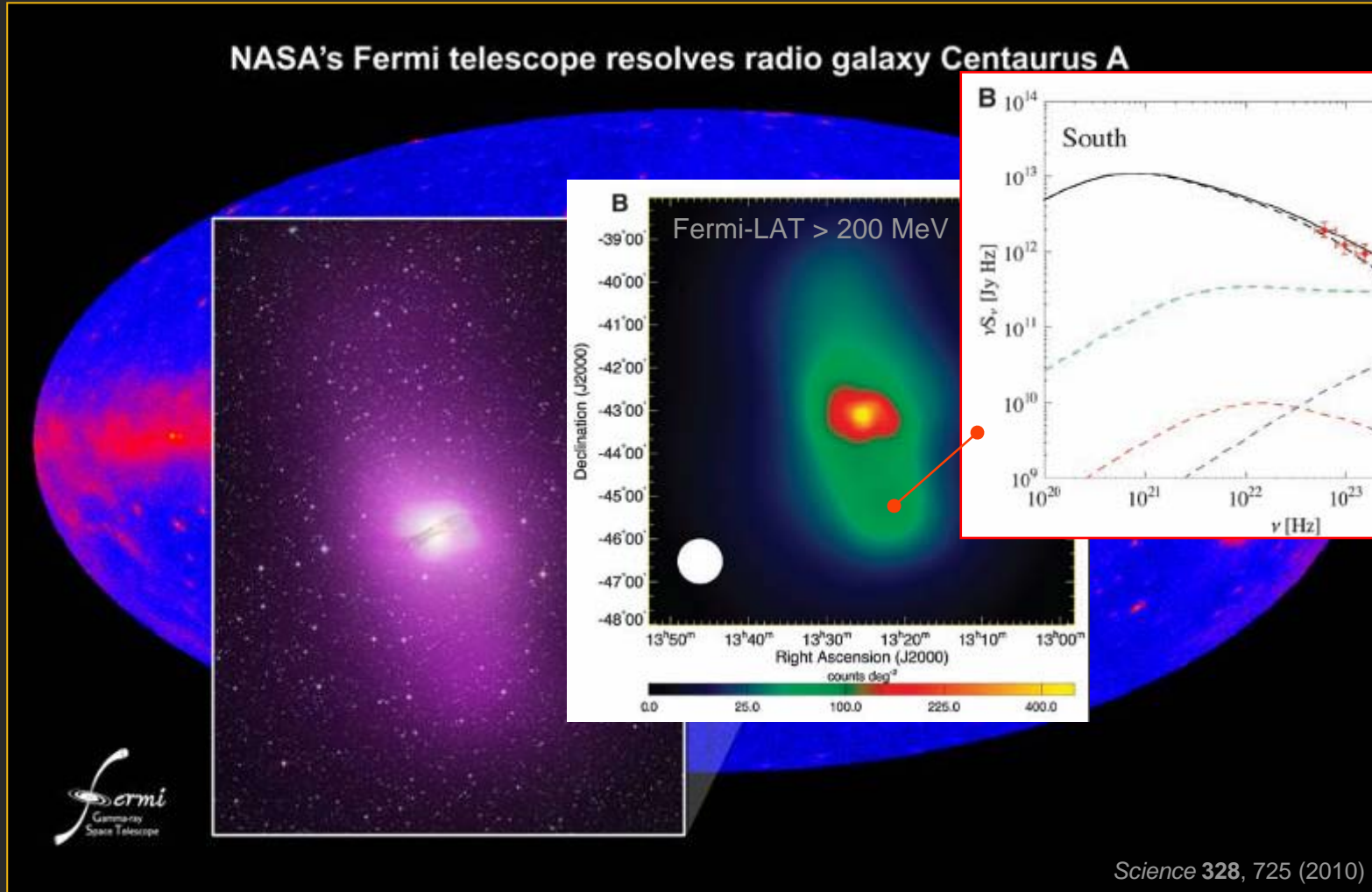
New category of Gamma-sources

- ◆ HE γ -ray emission from **starburst galaxies**
 - from interaction of CRs with local interstellar gas and radiation fields
 - depends on SN rate and large quantities of target gas

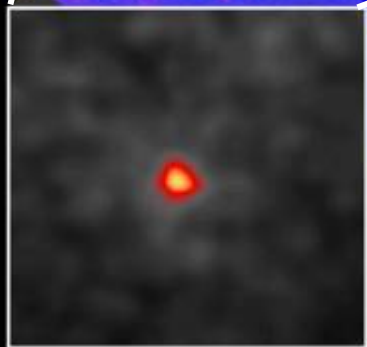
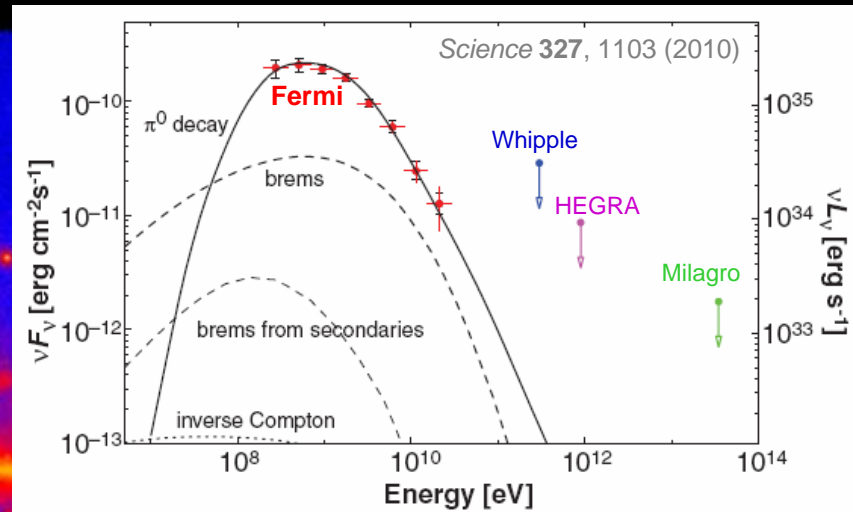
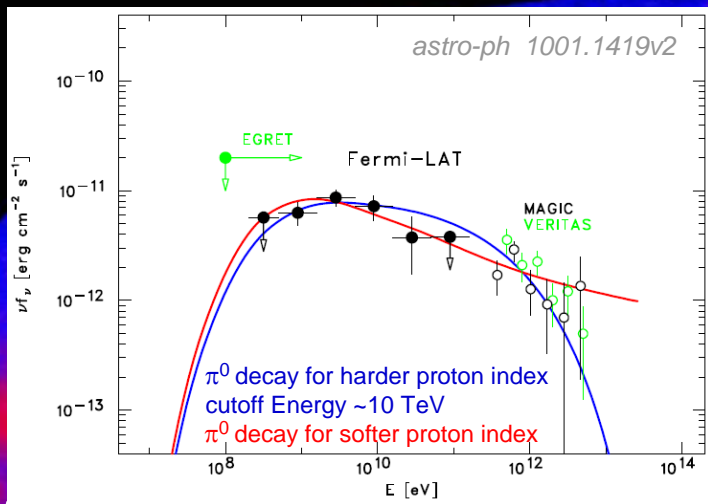


Imaging of a Radio Galaxy: Centaurus A

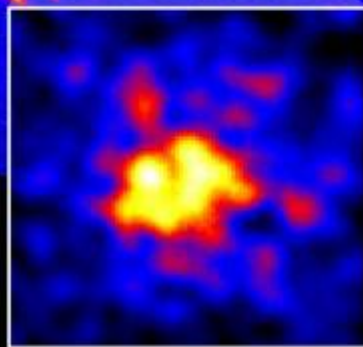
- ◆ Gamma-ray emission from the lobes:
 - Inverse Compton-scattered relic radiation from the CMB



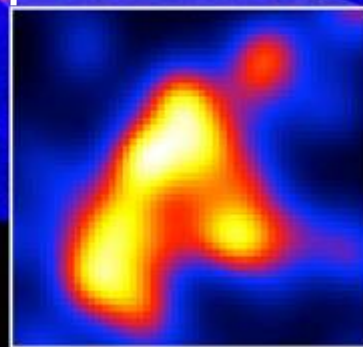
NASA's Fermi telescope resolves supernova remnants at GeV energies



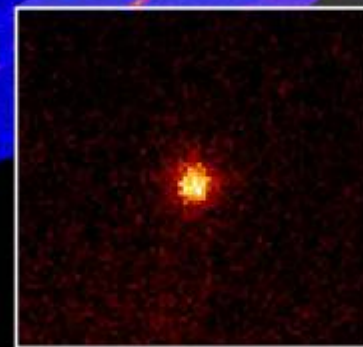
Cas A
young SNR



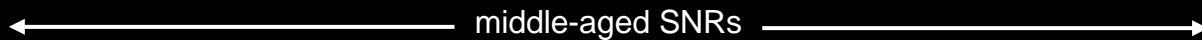
W51C



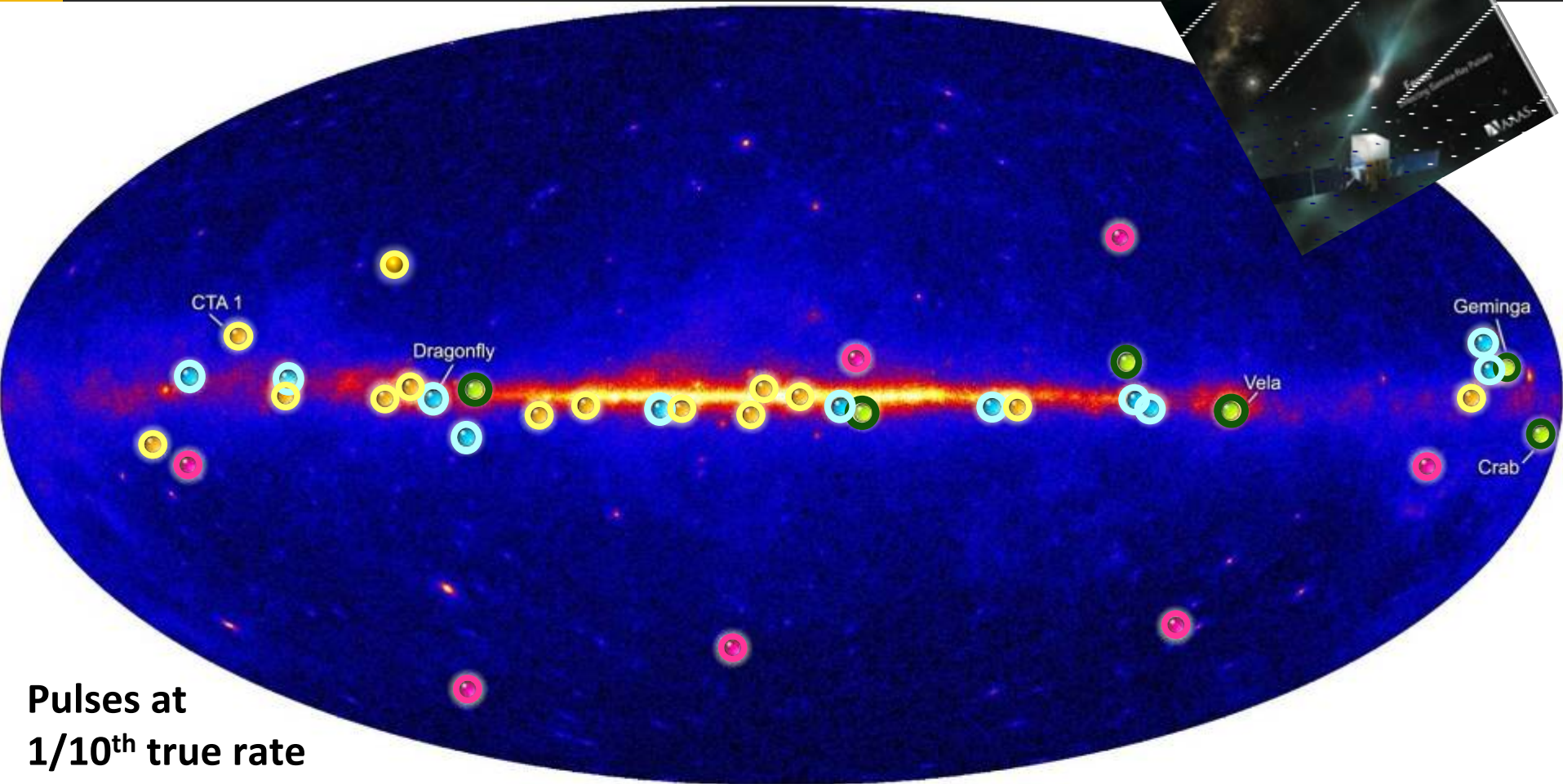
W44



IC 443



6-Month Pulsar catalog: the blinking Gamma-ray sky



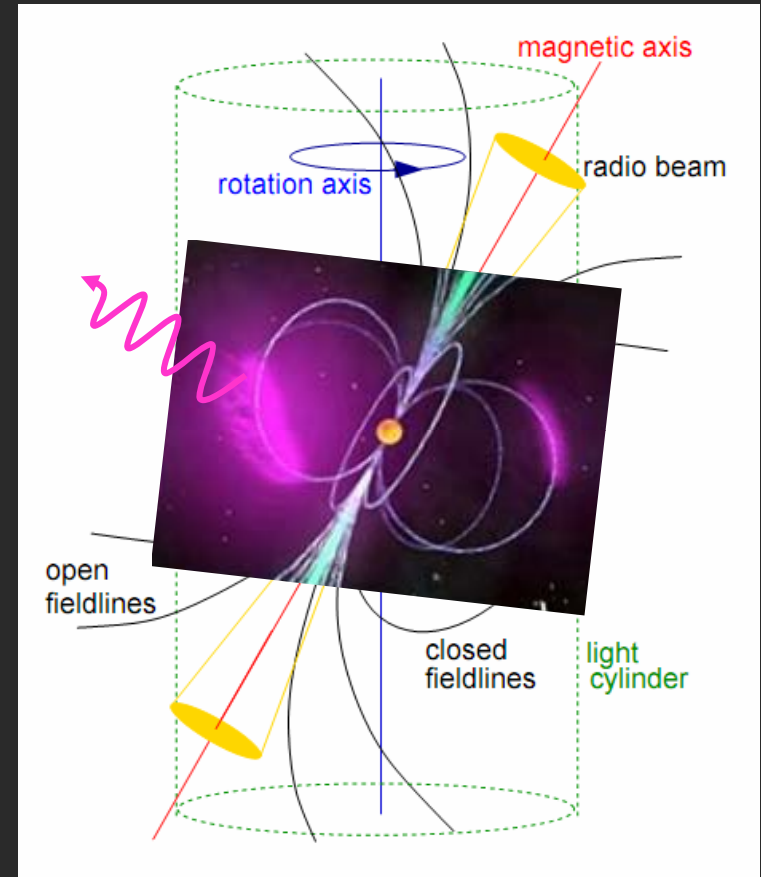
Pulses at
1/10th true rate

Fermi Pulsar Detections

- New pulsars discovered in a blind search
- Millisecond radio pulsars
- Young radio pulsars
- Pulsars seen by Compton Observatory EGRET instrument

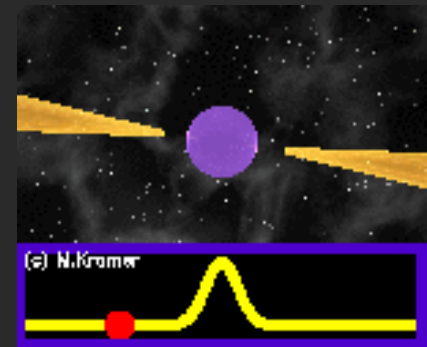
High-energy emission of Pulsars

- ◆ Fermi detects 16 new Pulsars
 - GeV photons are produced with a very high efficiency and into 'wide-angle' beams
 - Gamma-ray emission beam is much wider than radio emission
 - Most (but perhaps not all) observed light-curves can be reproduced in the outer-gap / TPC model



Andreas von Kienlin

E. Kendziorra's 65th birthday Workshop - IAAT

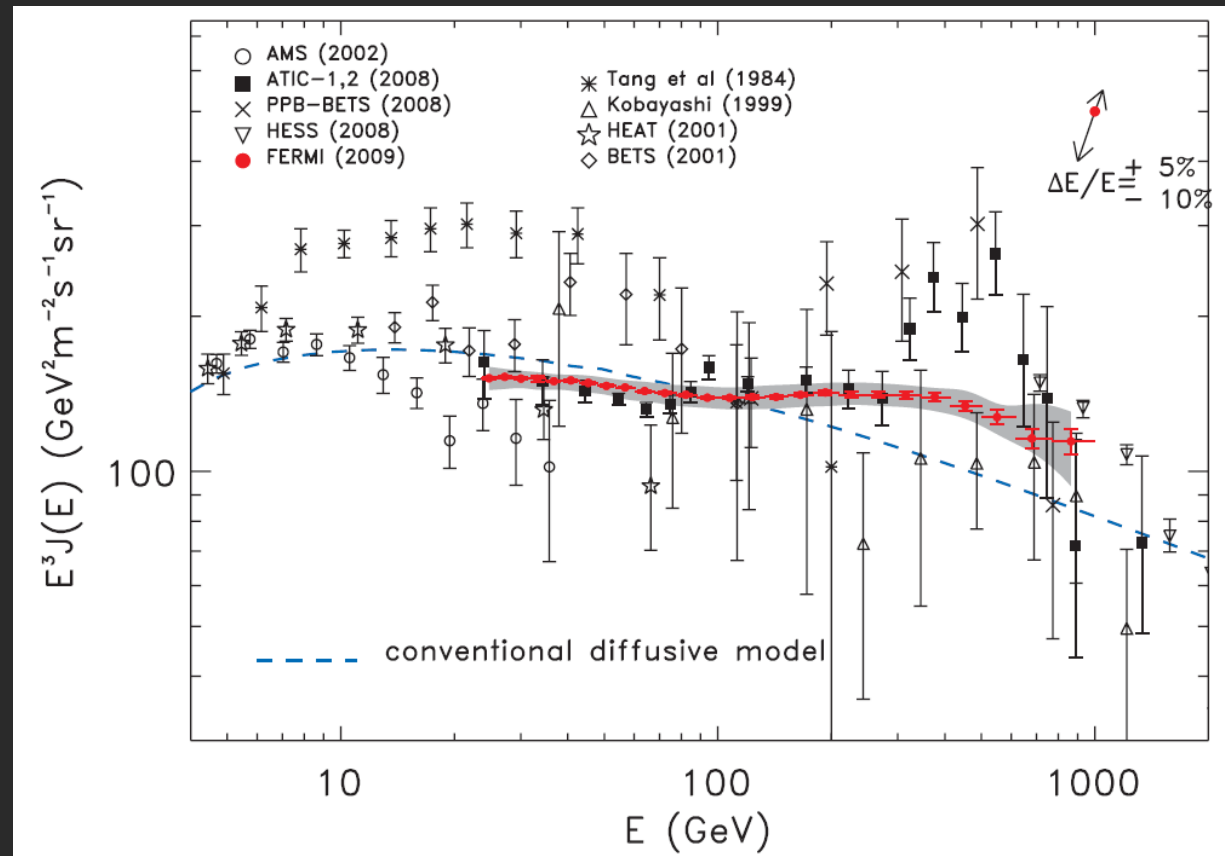


Cosmic Ray Electron Spectrum

- ◆ Fermi measures cosmic ray $e^- + e^+$ with $>2 \text{ m}^2\text{sr}$ acceptance (need to reject CR proton background)
- ◆ CR electron spectrum $E^{-3.04}$ between ~ 25 and 900 GeV
 - featureless; consistent with power law

- ◆ harder than predicted by GALPROP
- ◆ Local sources?
- ◆ Dark matter?

PRL 102, 181101 (2009)

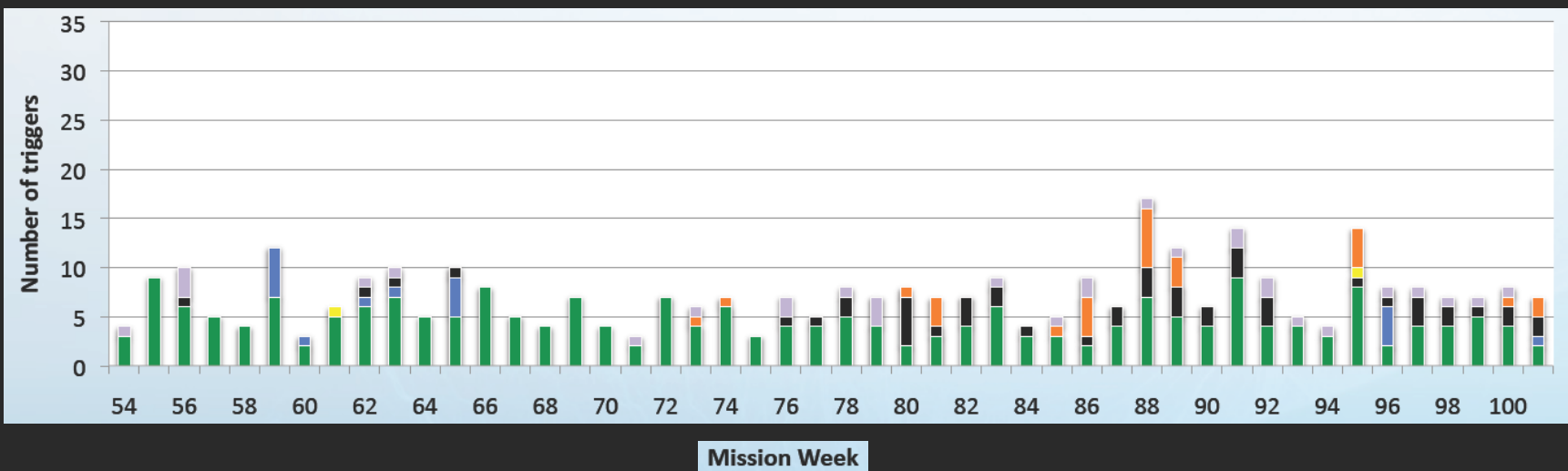
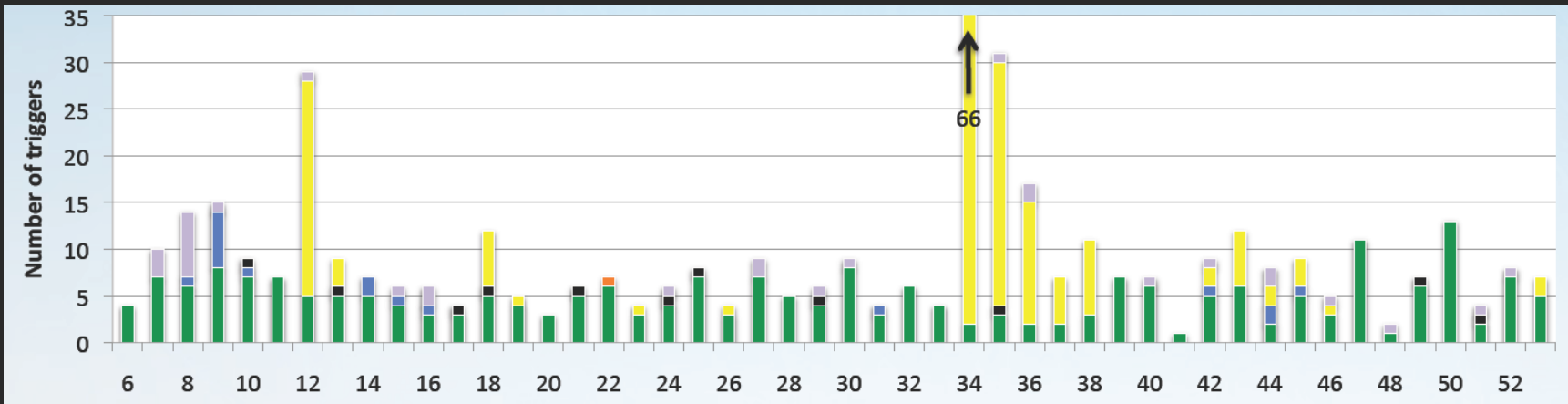
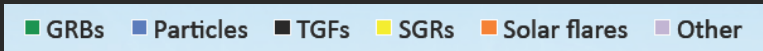


Gamma-ray Burst Monitor Science Highlights

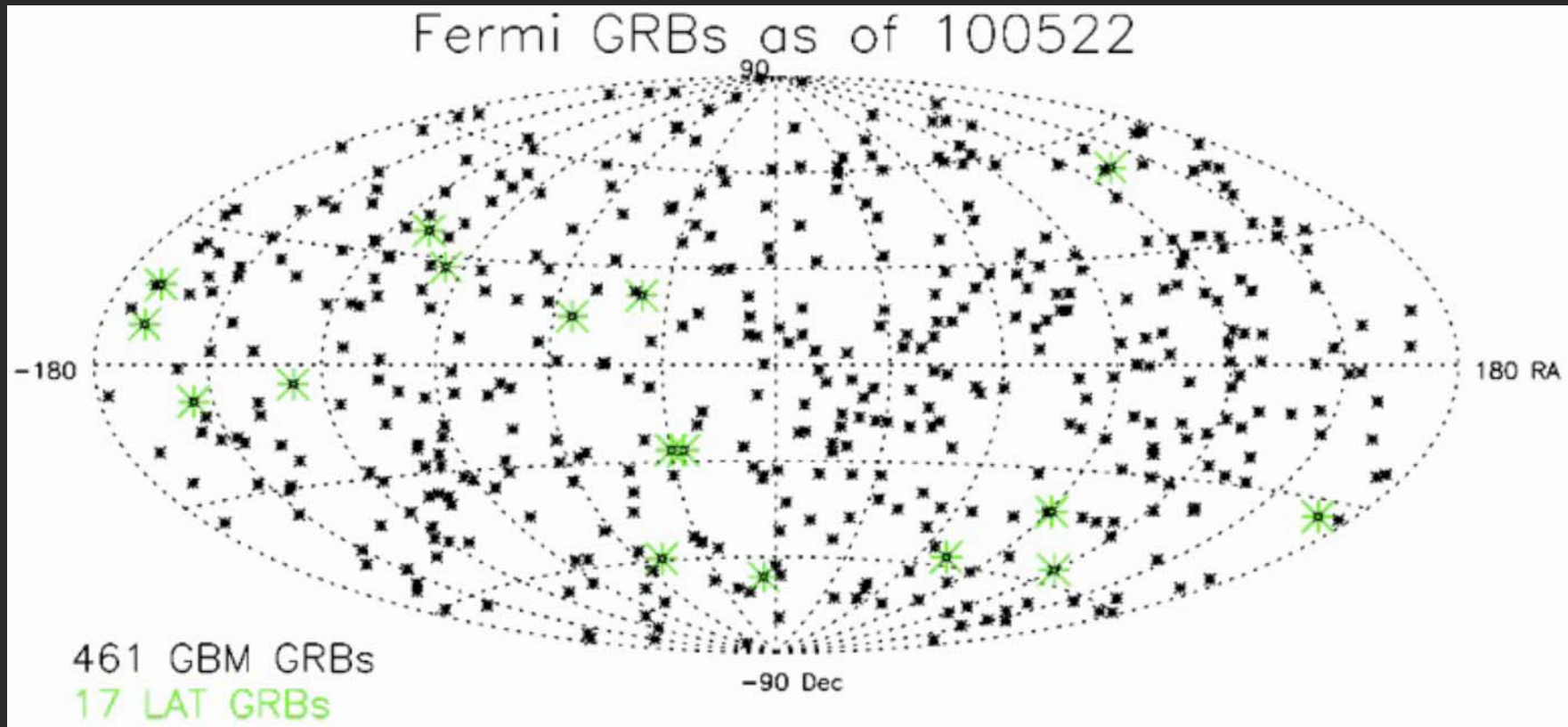
- ◆ Gamma-Ray Bursts (GRBs)
- ◆ Soft Gamma Repeaters (SGRs)
- ◆ Terrestrial Gamma-ray Flashes (TGFs)
- ◆ Solar Flares (SFs)



GBM Trigger History



Fermi GRBs

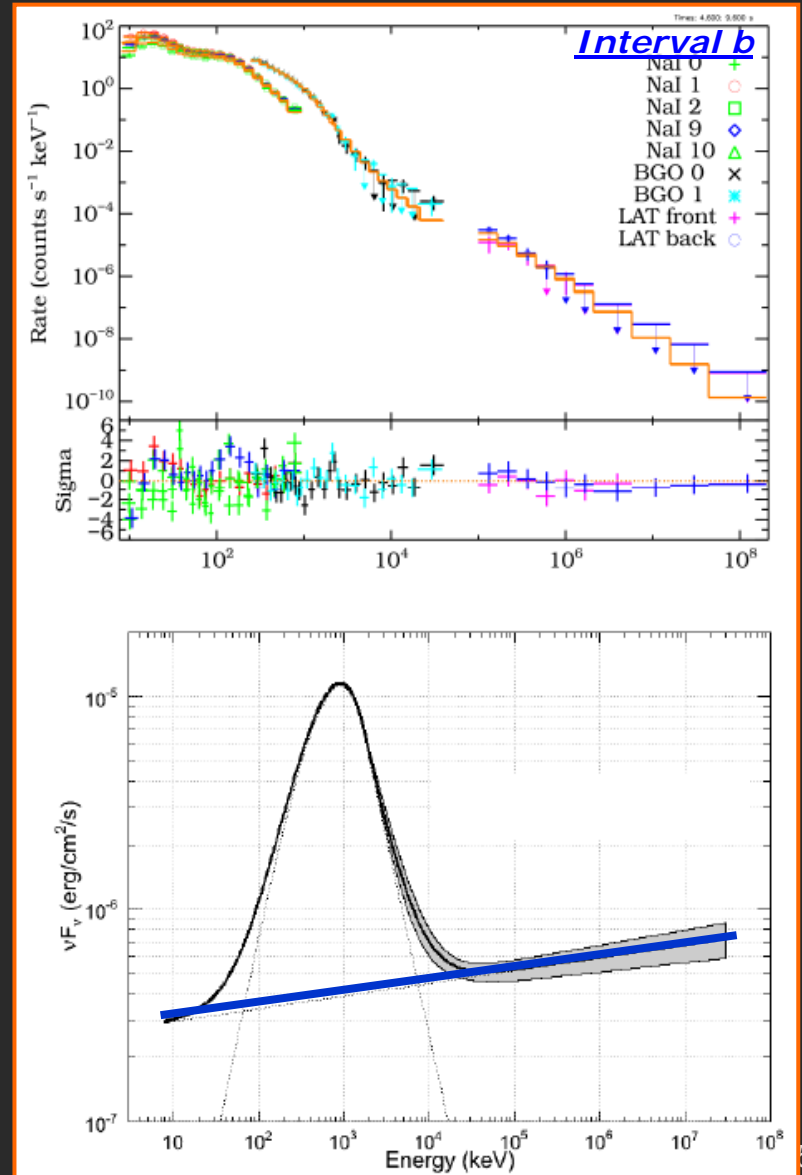
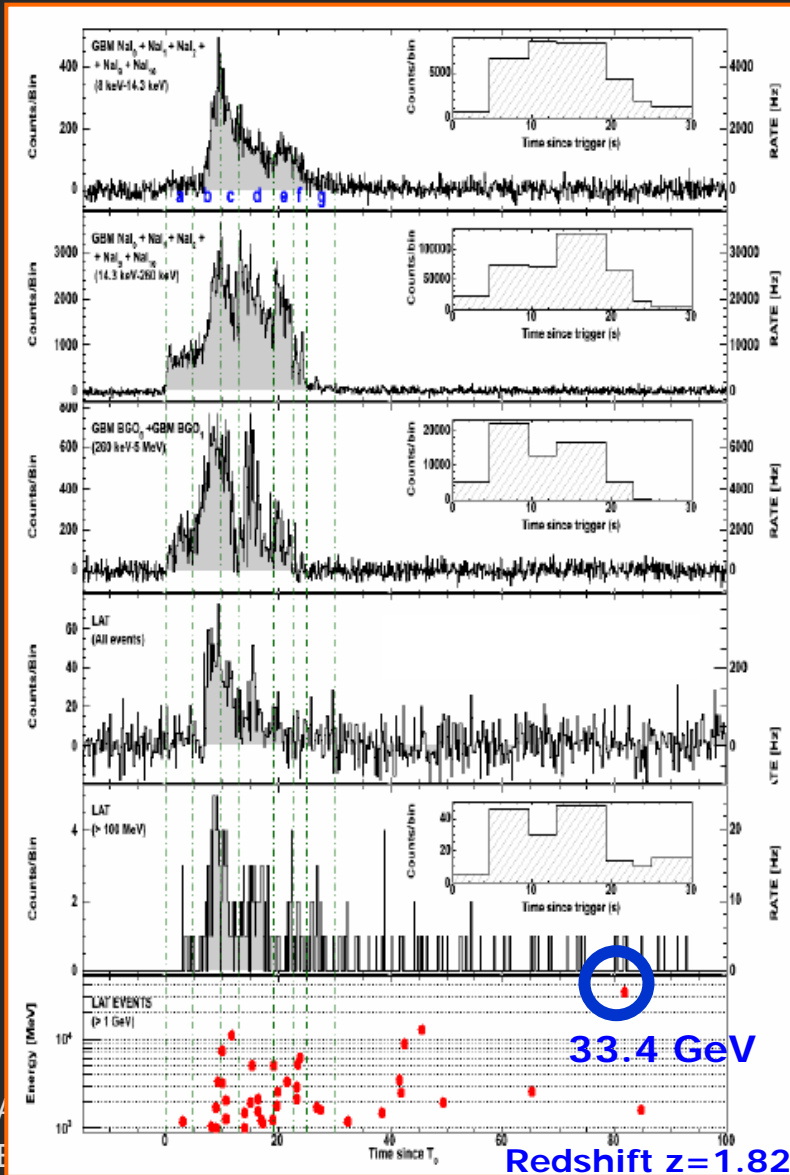


- ◆ 10 of 17 LAT bursts were observed (followed up) with Swift XRT, and all but 2 of these were detected by XRT.
- ◆ All XRT-detected LAT bursts were seen to have an optical counterpart, with a redshift measured in each case.

Fermi Observations of GRB 090902B

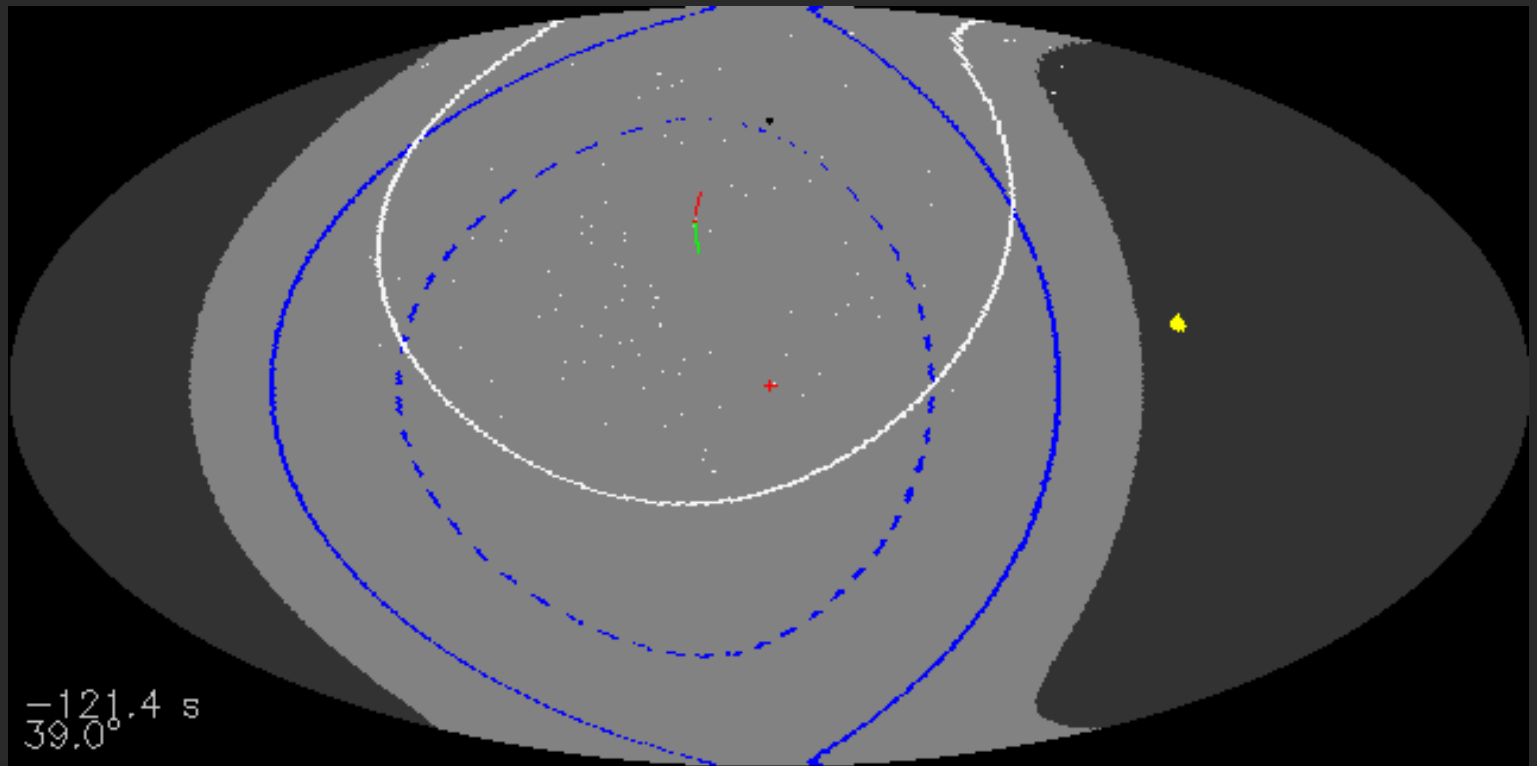
“A distinct spectral Component in the Prompt and Delayed emission”

Bissaldi, dePalma, Chiang & McBreen on behalf of the GBM and LAT Collaborations, ApJL submitted



Autonomous Repoint Request (ARR)

- ◆ ARR for GRB 090902B



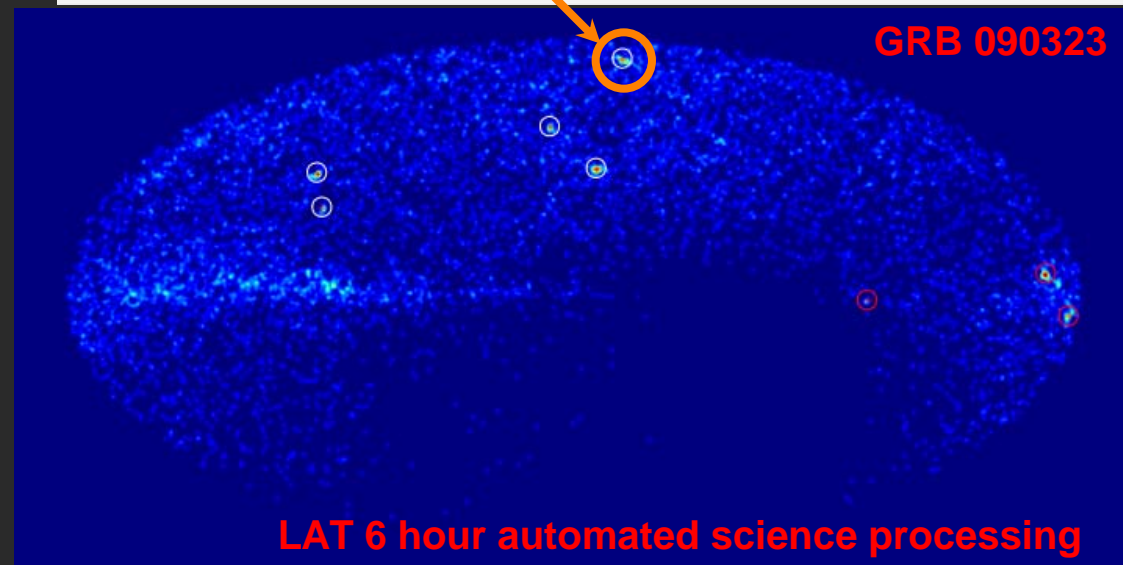
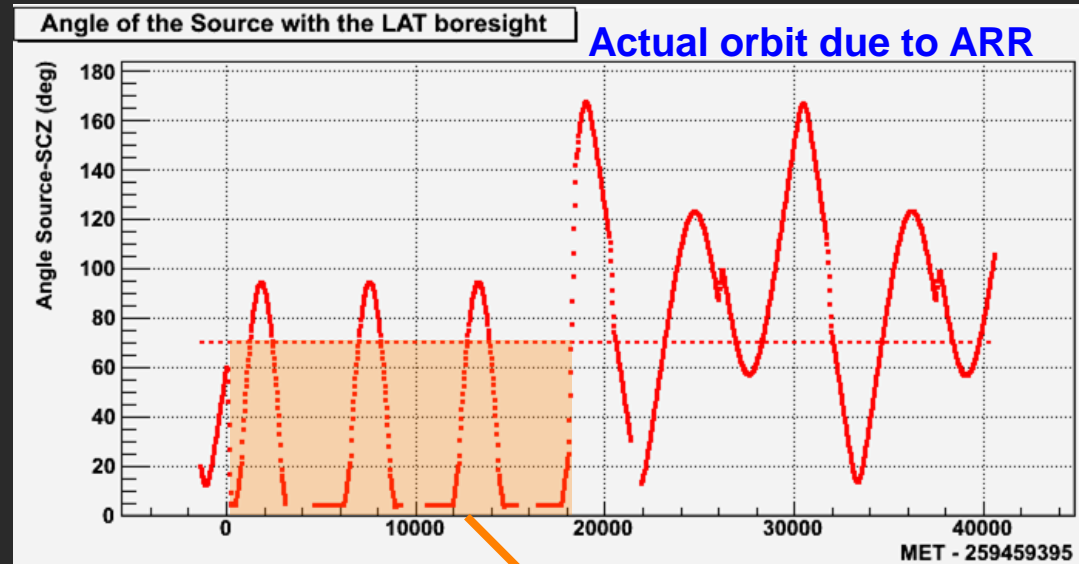
Long-Lived HE Emission in Other LAT GRBs

◆ GRB 090323

- HE emission up to a few GeV lasted ~2000 s.
- Made possible by ARR (autonomous repoint recommendation) by GBM
 - ▶ Otherwise GRB would have been out of FOV

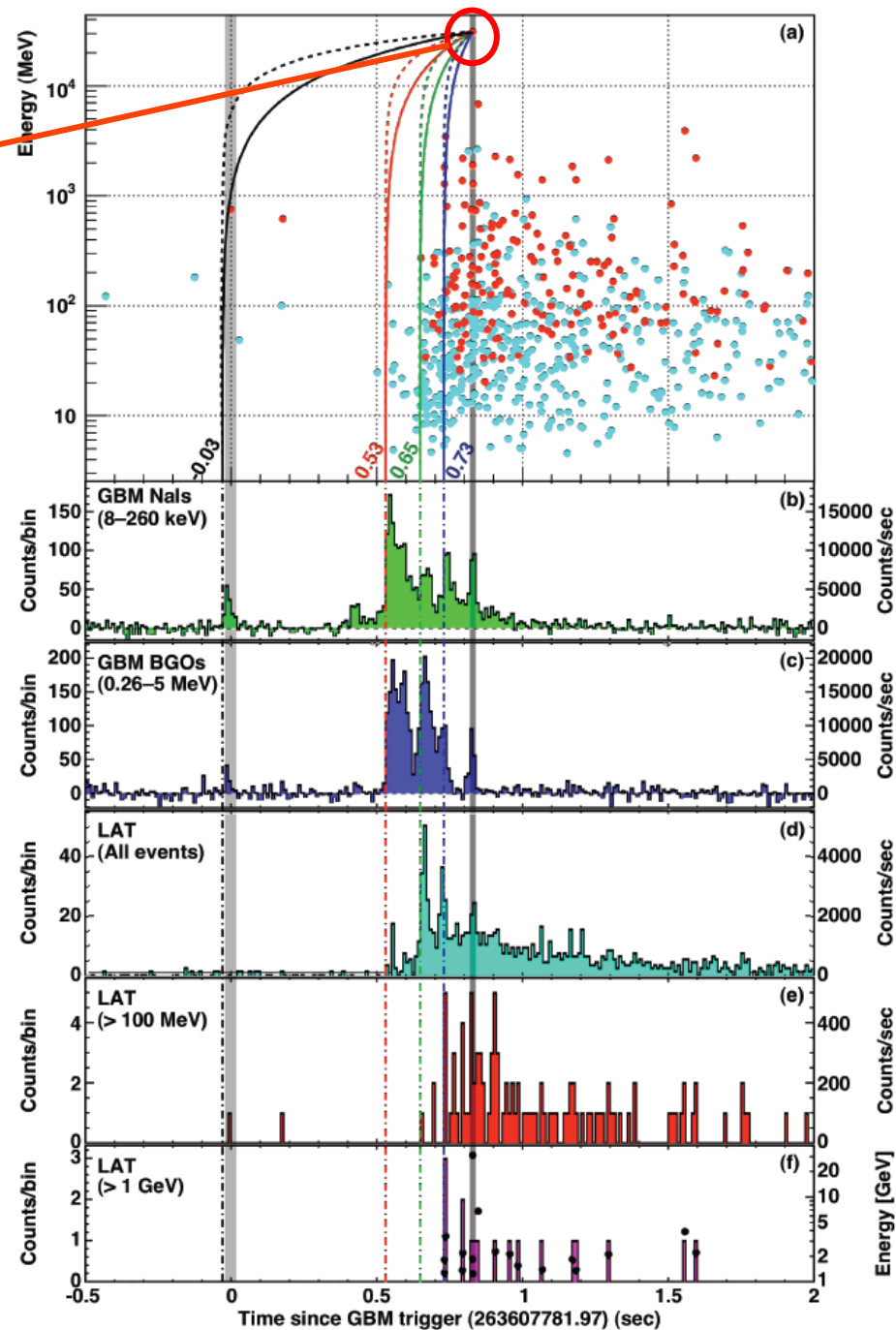
◆ GRB 090328

- HE emission lasted ~900 s
- ARR was also issued



GRB 090510

- ◆ 31 GeV photon: 0.83 s after the trigger
- ◆ This is the highest energy observed from short GRB
 - $z = 0.9 / 7.3$ billion ly
- ◆ Thus, this photon can be used to constrain both the bulk Lorentz factor of the relativistic jet and Lorentz Invariance Violation (LIV)
- ◆ GRB 080916C:
 - $M_{\text{QG}} > 0.1 M_{\text{Planck}}$
 - where $M_{\text{Planck}} \sim 10^{19}$ GeV
- ◆ GRB 090510 (short-hard burst):
 - $M_{\text{QG}} > \text{a few } M_{\text{Planck}}$
- ◆ At the fundamental length scale (Planck-scale of 1.62×10^{-33} cm) one expects that quantum effects affect nature of Space/Time



SGRs & AXPs in the *Fermi* era



SGR Source	Active Period	Triggers	Comments
J0501+4516	22 Aug - 03 Sep 08	26	New source at Perseus arm
1806-20	29 Nov 2008	1	Old source - reactivation
J1550-5418	03 - 20 Oct 08 22 Jan - 24 Feb 09 22 Mar - 17 Apr 09	7 117 14	Known source - first time exhibiting burst active episodes
J0418+5729	05 Jun 2009	2	discovered with GBM ! new source at Perseus arm
1833-0832	19 Mar 2010		discovered with Swift and RXTE - no GBM detection

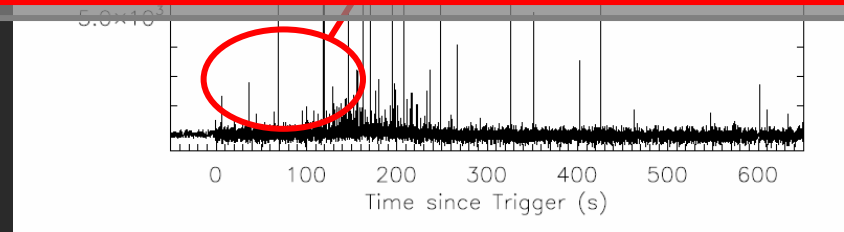
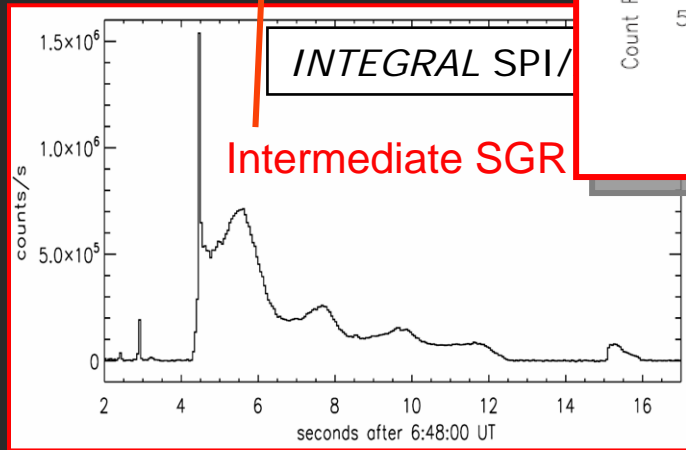
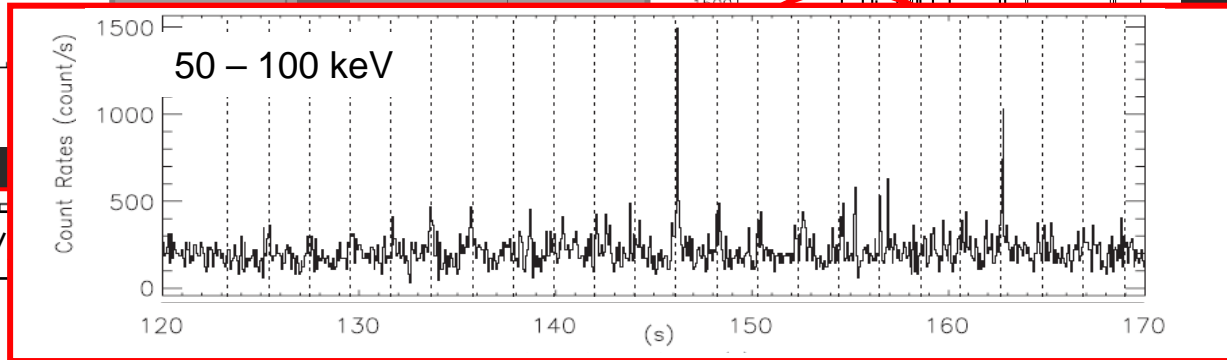
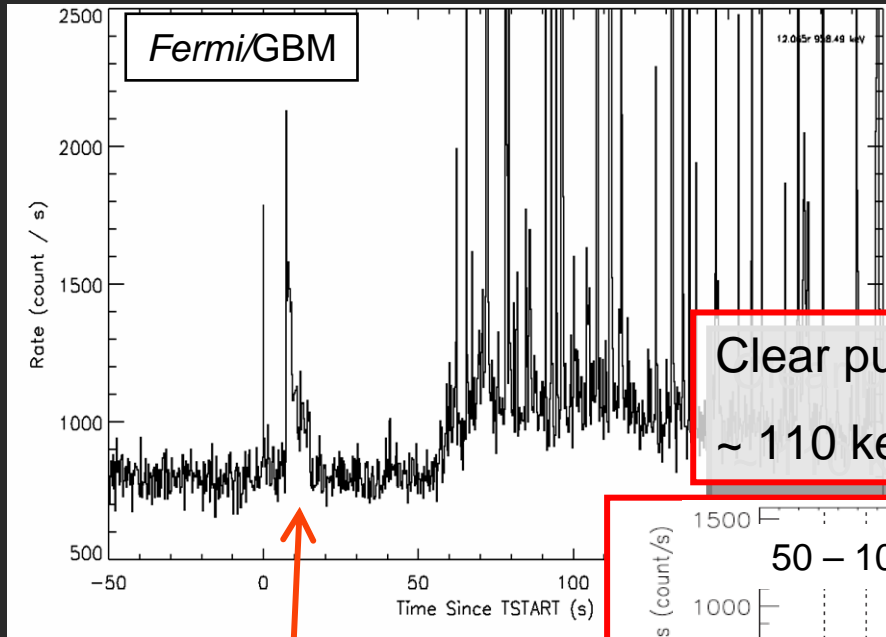
- ◆ GBM Magnetar Key Project (PI: Chryssa Kouveliotou)
 - <http://gammaray.nsstc.nasa.gov/gbm/science/magnetars>

Observations of the Magnetar SGR J1550-5418 with *Fermi* /GBM and *INTEGRAL* SPI/ACS

- ◆ 3 activity periods: Oct 08, Jan 09, March & April 09
 - About 500 bursts detected
 - Two intermediate flares
 - Enhanced persistent emission with pulsations at the NS spin period

Clear pulsation up to
~ 110 keV seen!

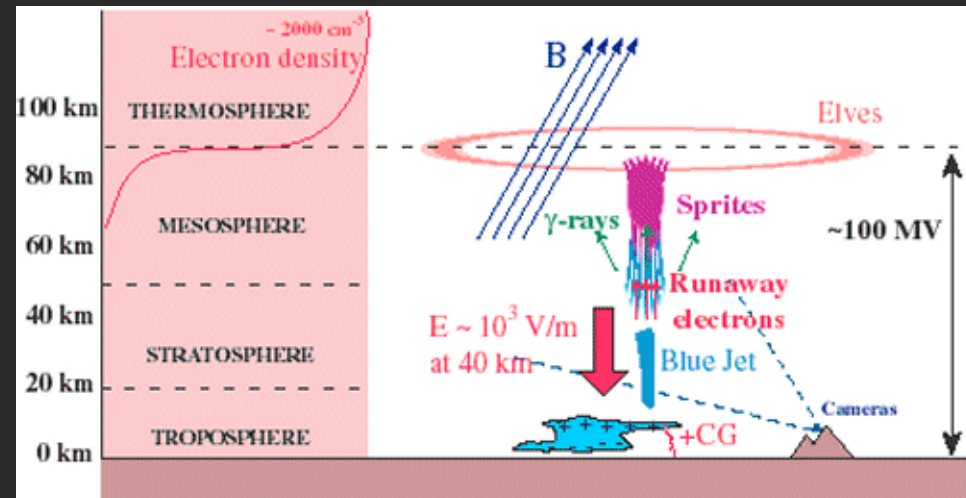
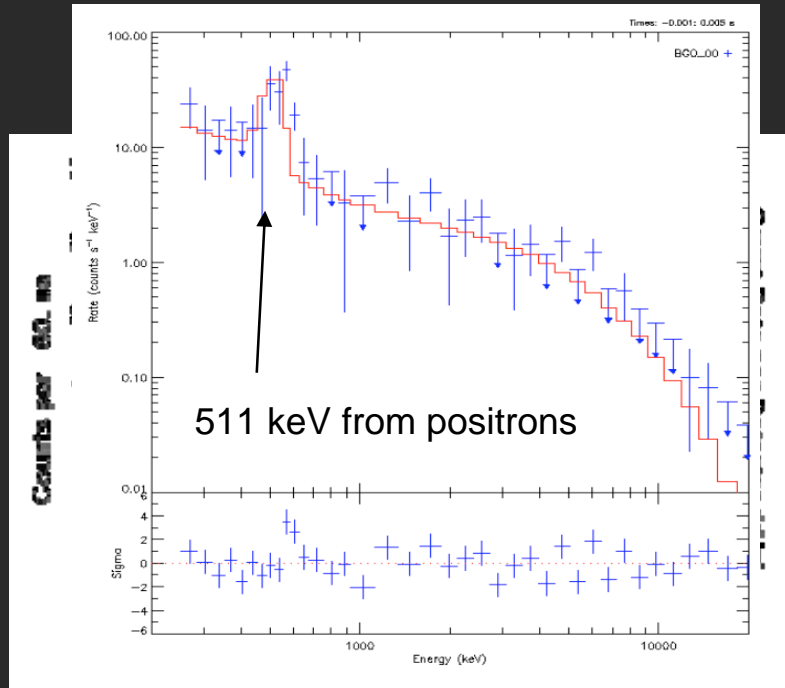
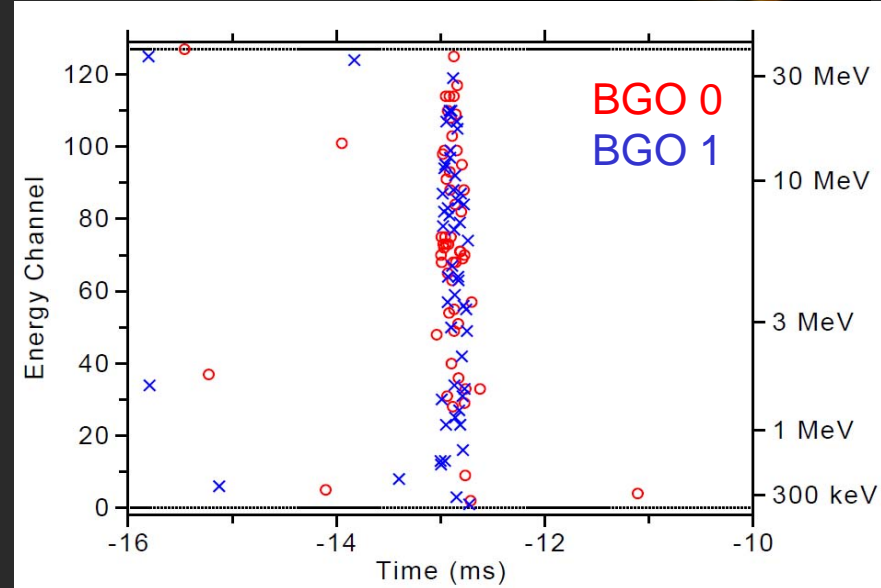
~ 150 s long enhanced persistent Emission



Terrestrial Gamma-ray Flashes (TGFs)

Fermi

- ◆ GBM triggered on 79 TGFs
- ◆ Two types of TGFs:
 - Short (~ 0.1 ms) γ -ray TGFs
 - Long (~ 1 ms) electron/positron TGFs



Andreas von Kienlin

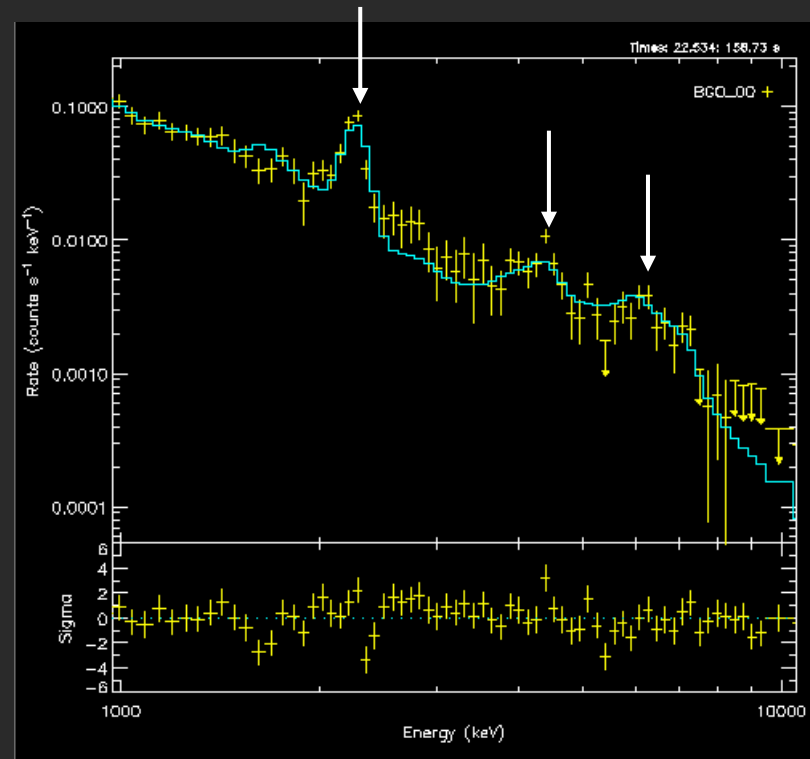
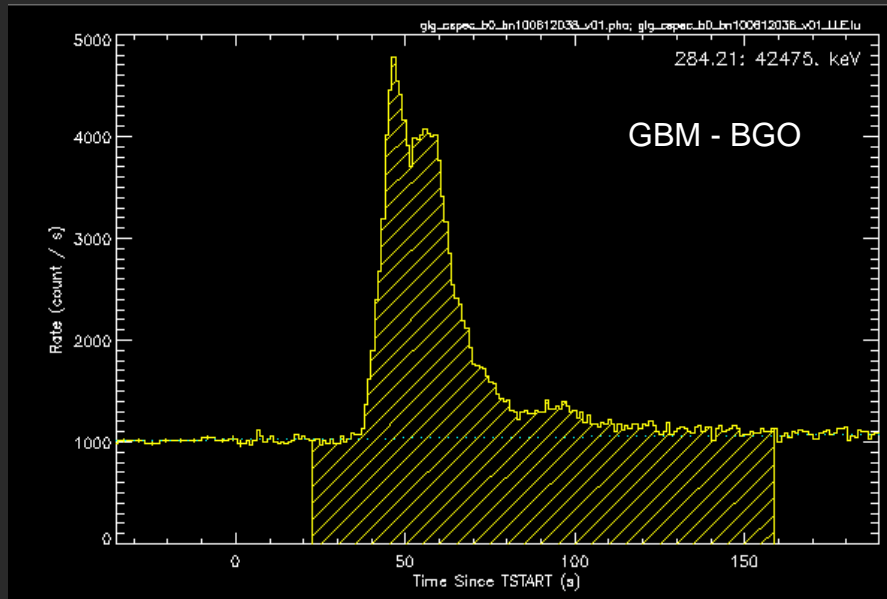
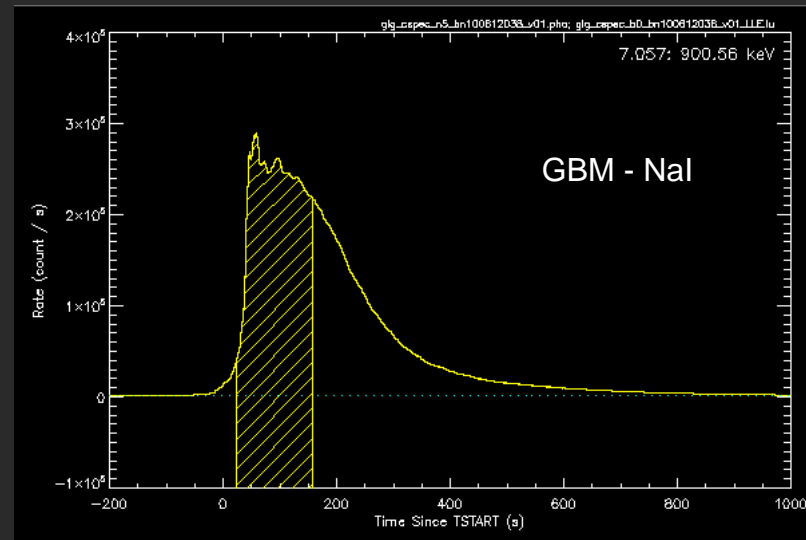
E. Kendziorra's 65th birthday Workshop - IAAT

July 15, 2010

31

Solar Flares (GBM)

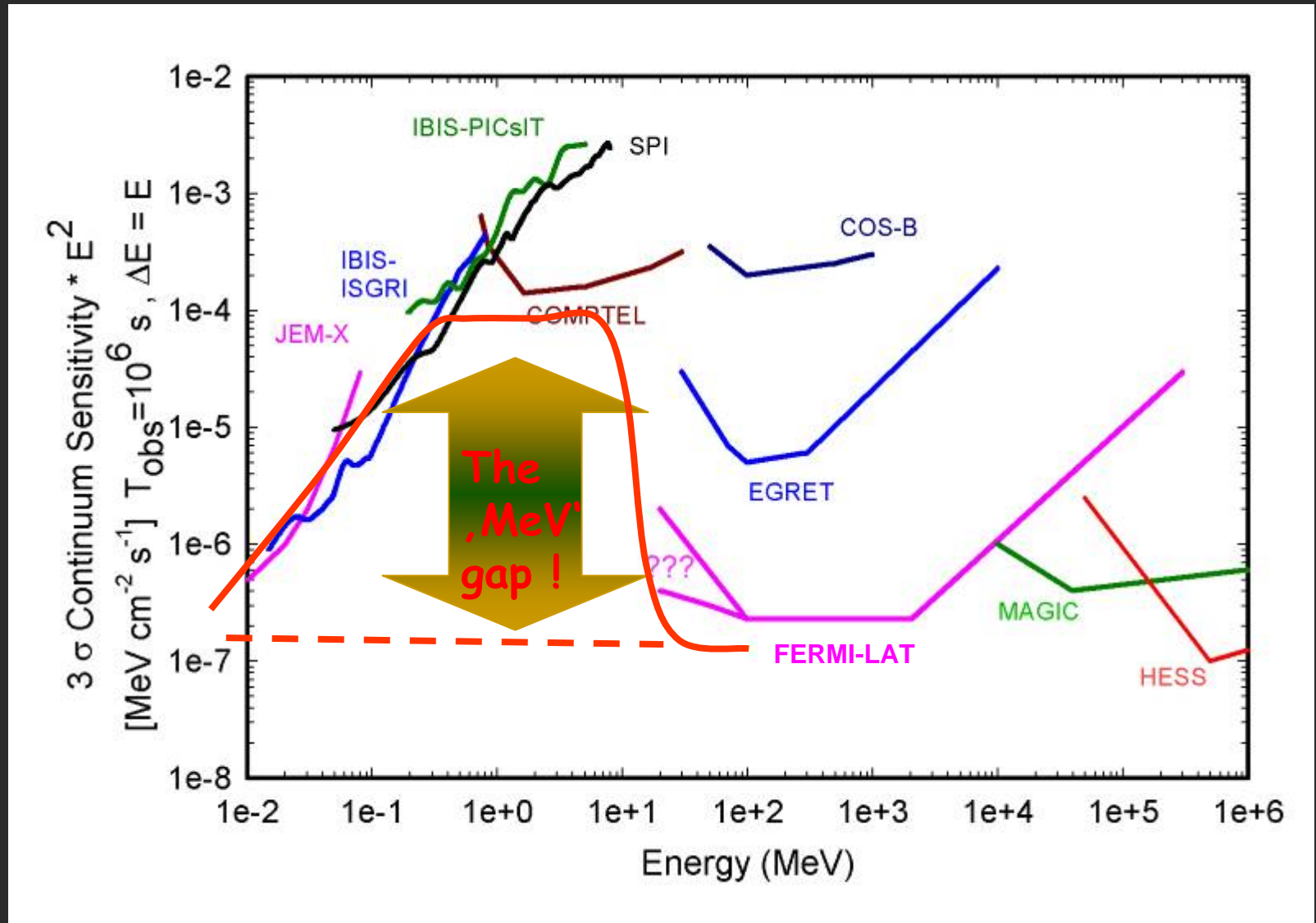
- ◆ June 12, 2010: First solar flare with nuclear lines
 - 2.2 MeV ^1H neutron capture line
 - 4.4 MeV $^{12}\text{C}^*$ nuclear interaction line
 - 6.1 MeV $^{16}\text{O}^*$ nuclear interaction line



Highlights Summary

- ◆ Large Area Telescope (*Fermi-LAT*):
 - 1st Fermi LAT Source catalog (1451 sources)
 - Observes flaring and variable source
 - Starburst galaxies + LMC detected in GeV γ -rays
 - Multiwavelength Campaigns
 - Resolves SNRs at GeV energies
 - γ -ray emission from Cen A's radio lobes
 - LAT detects many pulsars and constrains emission mechanism
- ◆ Gamma-ray Burst Monitor (*Fermi-GBM*)
 - 500 GRBs in 2 years
 - 17 LAT GRBs \Rightarrow extra emission component, delayed onset, LAT/GBM constrain LIV
 - Burst from 4 SGRs detected
 - TGFs
 - 1st Solar Flare with nuclear lines

Outlook: Future Mission in the Gamma-ray regime ?



Outlook: Future Mission in the Gamma-ray regime ?

- ◆ Advanced Compton Telescope (ACT)
- ◆ GRIPS

